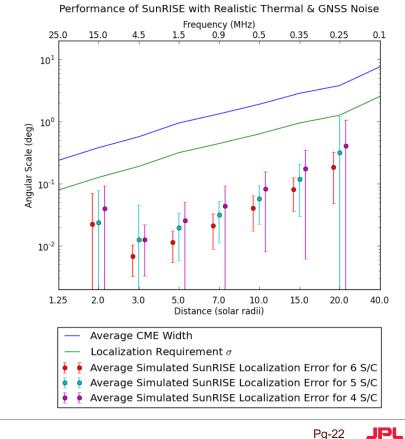


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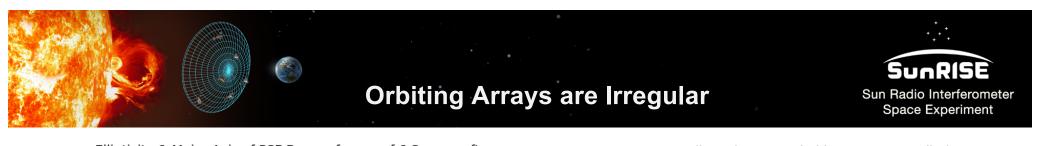
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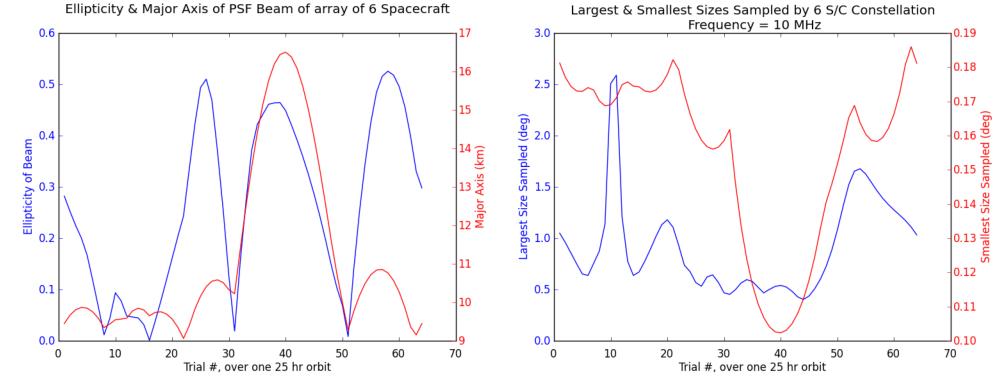


- Localization resolution determined by array configuration and frequency.
- SunRISE localization resolution ranges from 0.01 – 0.1 degrees (or 0.6 – 6 arcseconds)



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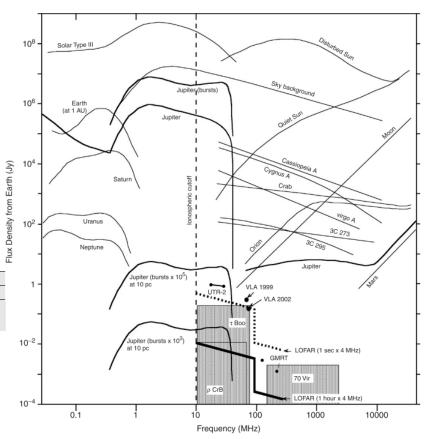


- SunRISE might be first array that can make basic images at low Frequencies below lonospheric cutoff
- New Window into the universe, plenty to look at
- Astronomy review papers have detailed requirements for every possible source

Торіс	Our Sections	Requirements					
		Frequency/MHz	Resolution $\vartheta$	Baselines/km	Expected signal <sup>a</sup>	N(Antennas)	$t_{\rm exp}~(5-\sigma)$
Transients	3.4						
Solar and planetary bursts	3.4.1	0.1-30	degrees	0.5-200	MJy	1-100	min-h
Extrasolar planets	3.4.2	0.5-30	$\lesssim 1'$	$\gtrsim 35{-}1000$	1–10 mJy	$10^4 - 10^{5f}$	15 min





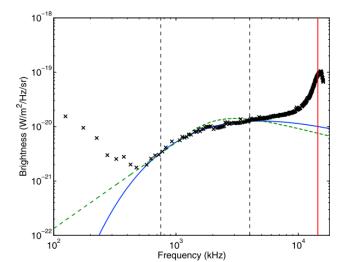




- Mirror galactic calibration of STEREO antenna from Zaslavsky et al 2011
- Must understand Antenna and Stray impedance, goes into  $\Gamma^2$

$$V_r^2 = V_{noise}^2 + \Gamma^2 V_{QTN}^2 + \frac{4\pi}{3} Z_0 \Gamma^2 l_{eff}^2 B_f.$$

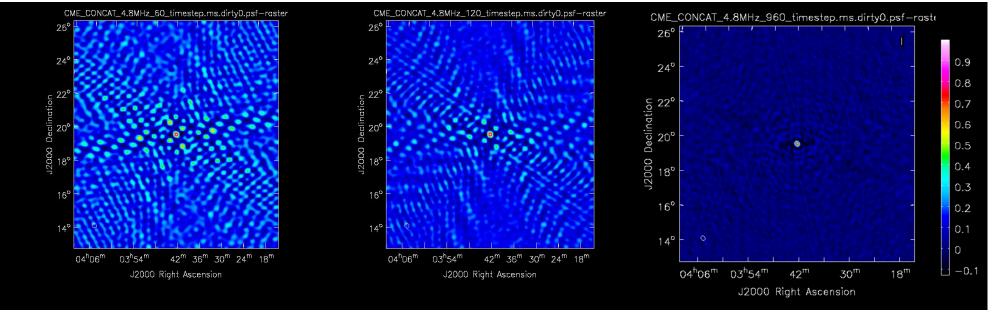
- Choose middle range where galactic noise is dominant (Quasi Thermal Plasma Noise dominates at lowest freqs, short antenna approx. fails at higher freqs)
- Subtract off constant antenna noise to solve for effective antenna length
- Compare calibrated data (crosses) with Galactic brightness models Novacco and Brown [1978] (blue solid line) and Cane [1979] (green dashed line)



Zaslavsky et al. RADIO SCIENCE, VOL. 46, RS2008, doi:10.1029/2010RS004464, 2011

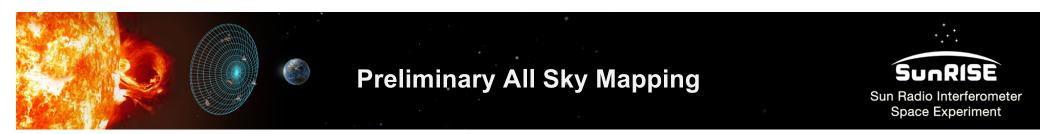




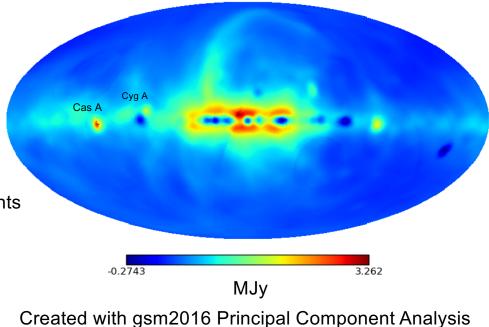


PSFs at 5 MHz after 1, 2 and 24 hours of integration. All images share the same color scale, normalized to 1.0.





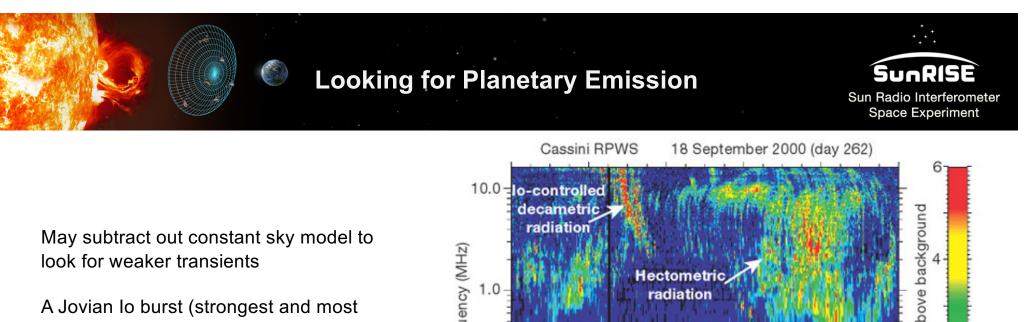
- Will have 1 year's worth of data to sift through
- Choose least bright periods to create constant Sky Maps
- For given area of sky, use SPICE to compute times when it is unobscured by Earth, Sun, Jupiter, etc.
- · Can subtract constant sky model to look for weak transients
- Working on simulated SunRISE Global Sky Map



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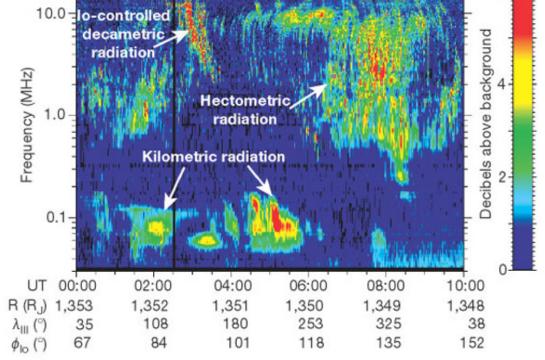
Zheng et al. MNRAS 464, 3486-3497 (2017)

20 MHz Galactic Sky Map Mollweide Projection



predictable) will typically dwell on a frequency for > 1 hour.

Processing of searching for Jovian Emission mirrors that of extrasolar planetary emission



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## **Jupiter's Io Decametric Radiation as a Calibration Source**

Property	Range	Notes		
Frequency	0.3 MHz – 35 MHz	Significant overlap with SunRISE band.		
Occurrence	- over couple of days	Predictable occurrence based on lo orbital		
	~every couple of days	phase and Jupiter's longitudinal phase.		
Duration	~ 2 hours	Equivalent ~72,000 snapshots with		
		SunRISE		
Flux Density	$10^{-20}$ to $10^{-19}$ W m <sup>-2</sup> Hz <sup>-1</sup>	Flux is variable but strong when active.		
		Stereo/Waves sees them regularly.		
Structure	Point source	Source size < 400 km at 4.4 AU from VLBI		
	Pollit Source	measurements.		

NOTE: These data are gathered while in science mode. It does not interfere with regular operations.

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- SunRISE, designed for Solar Radio Bursts, can see the entire Low Frequency Sky over 12 month mission
- SunRISE could make first maps of the Sky at these Frequencies
- Could do preliminary Galactic foreground subtraction
- SunRISE can localize individual radio sources
- Data Processing mirrors that of a larger array that could detect Extrasolar Planetary Emission
- Space Based Interferometry will be huge, SunRISE could be the pathfinder

