

Low Frequency Radio Emission

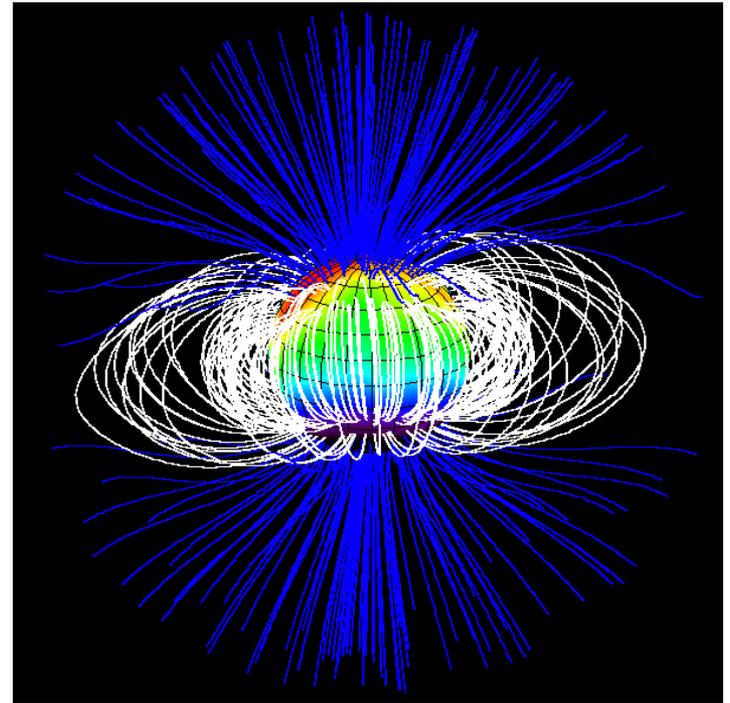
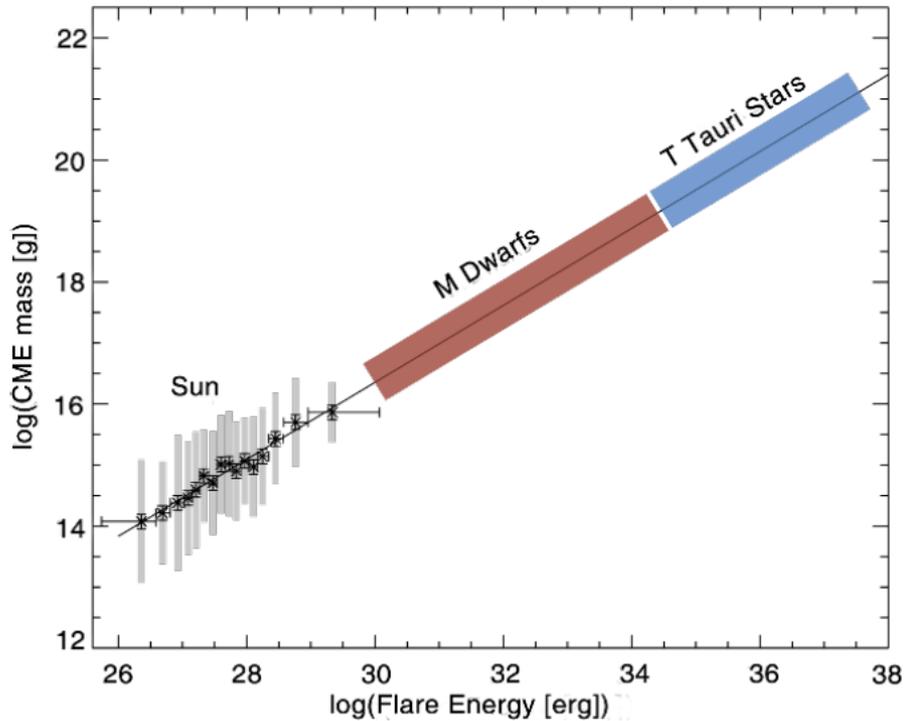


**Type II radio bursts
traces density at CME shock**

**Auroral radio emission
measures magnetic fields**



Stellar CMEs



Donati et al. 2006

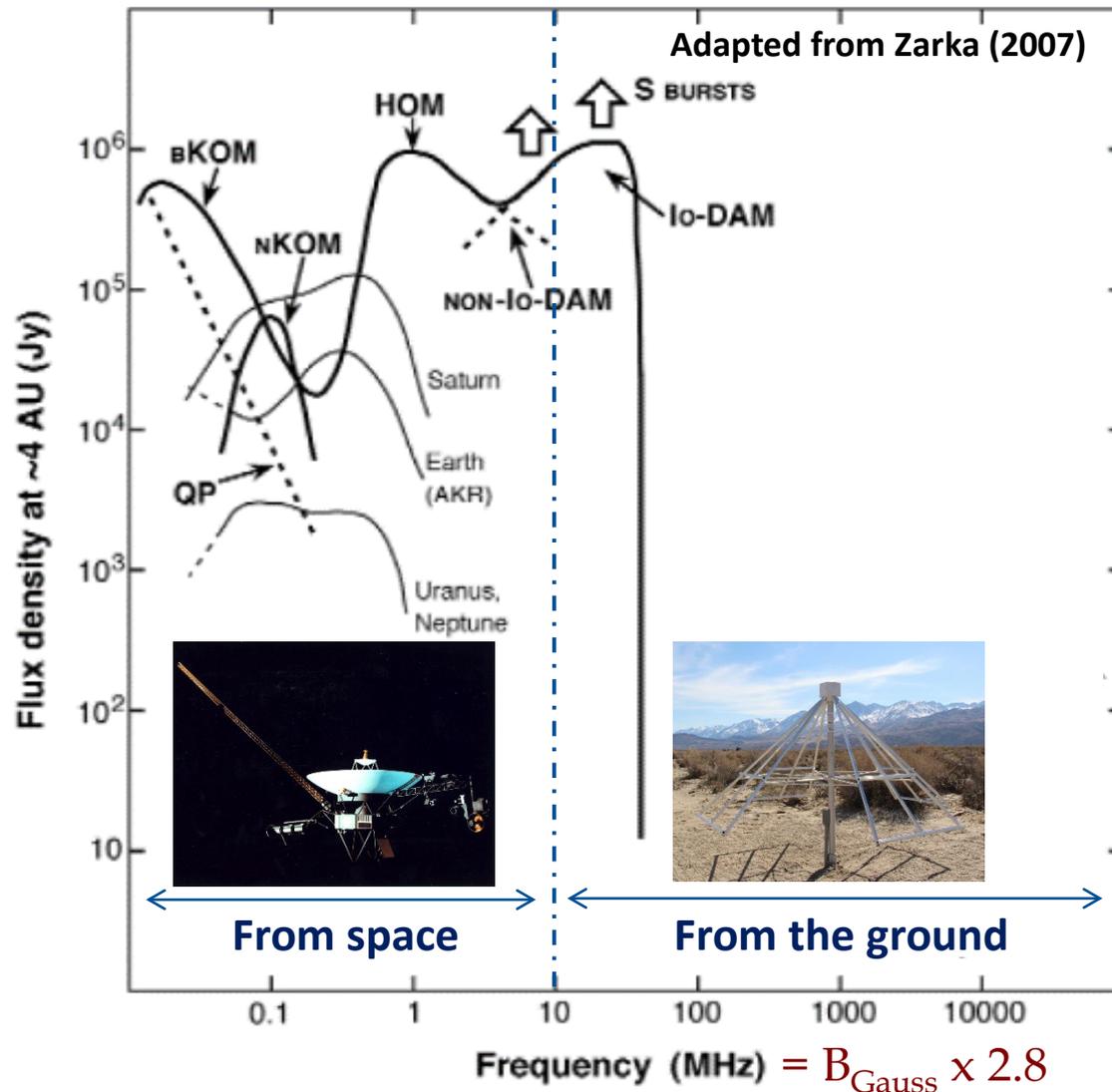
Adapted from Aarnio et al. 2012

No direct evidence of CMEs on any star other than the Sun to date

**Magnetic field configuration may be play an important role
(Alvarado-Gómez et al. 2018)**

Radio Emission from Solar System Planets

- ☐ All gas giants and Earth have strong auroral radio emission
- ☐ Electron cyclotron maser emission - coherent, highly circularly polarized



Paradigm Shift

Radio

Microwave

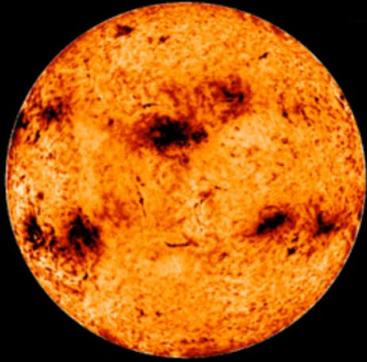
Infrared

UV

X-Ray

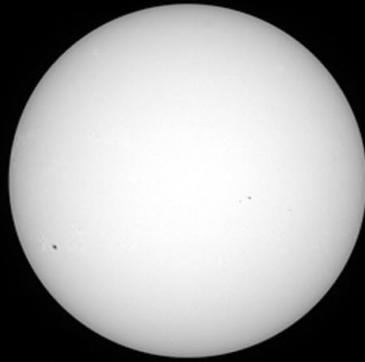
Gamma Ray

Sun: 1,083 nm



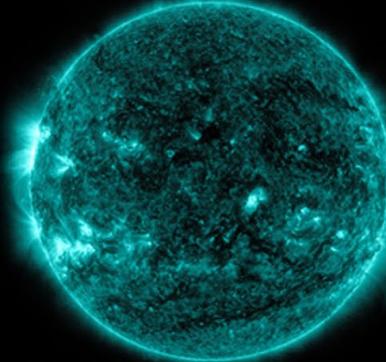
Infrared
Telescope

Sun: 750 – 380 nm



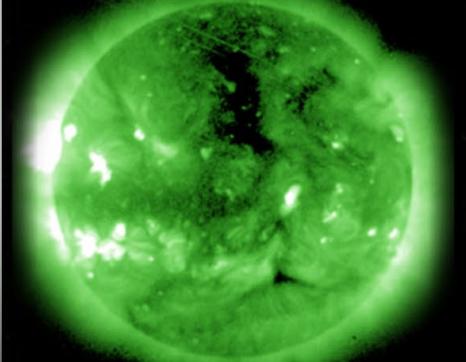
Optical
Telescope

Sun: 13.1 nm

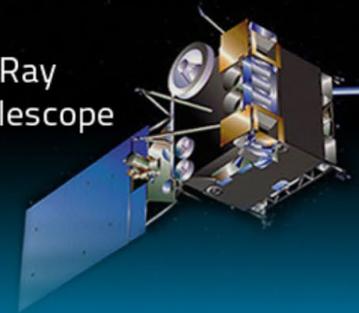
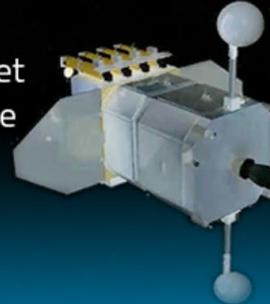


Ultraviolet
Telescope

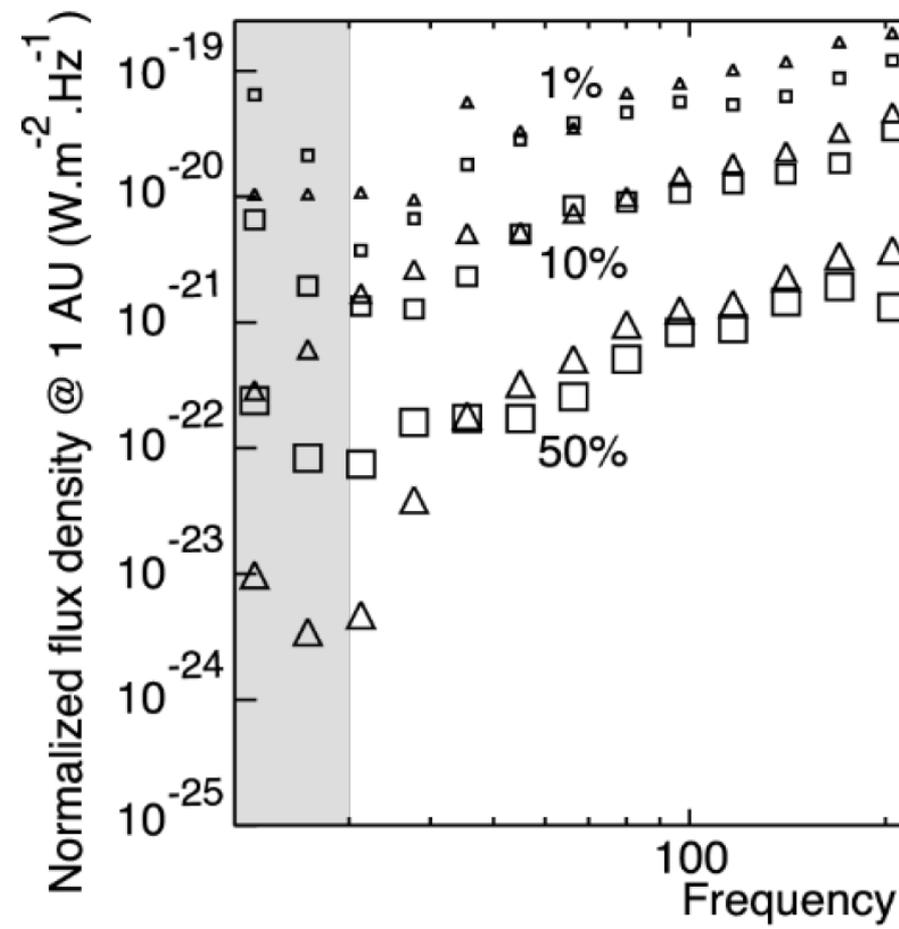
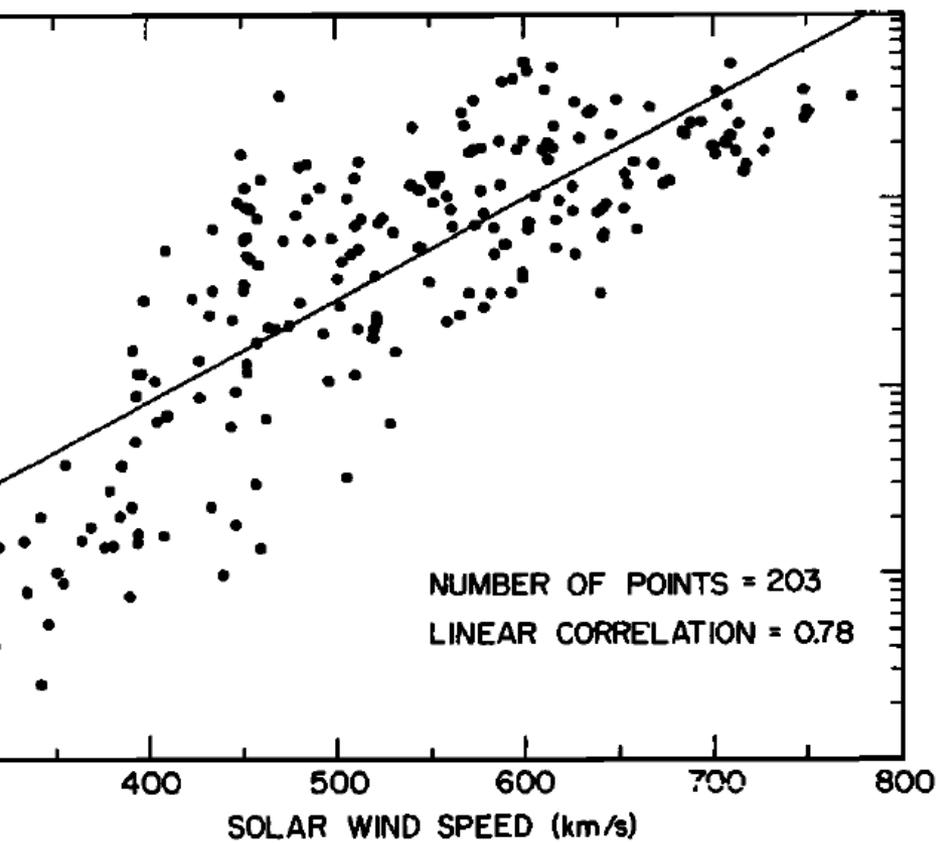
Sun: 6 – 0.6 nm



X-Ray
Telescope



SMOOTHED DAILY AVERAGES FOR DAYS 160-365 IN 1974

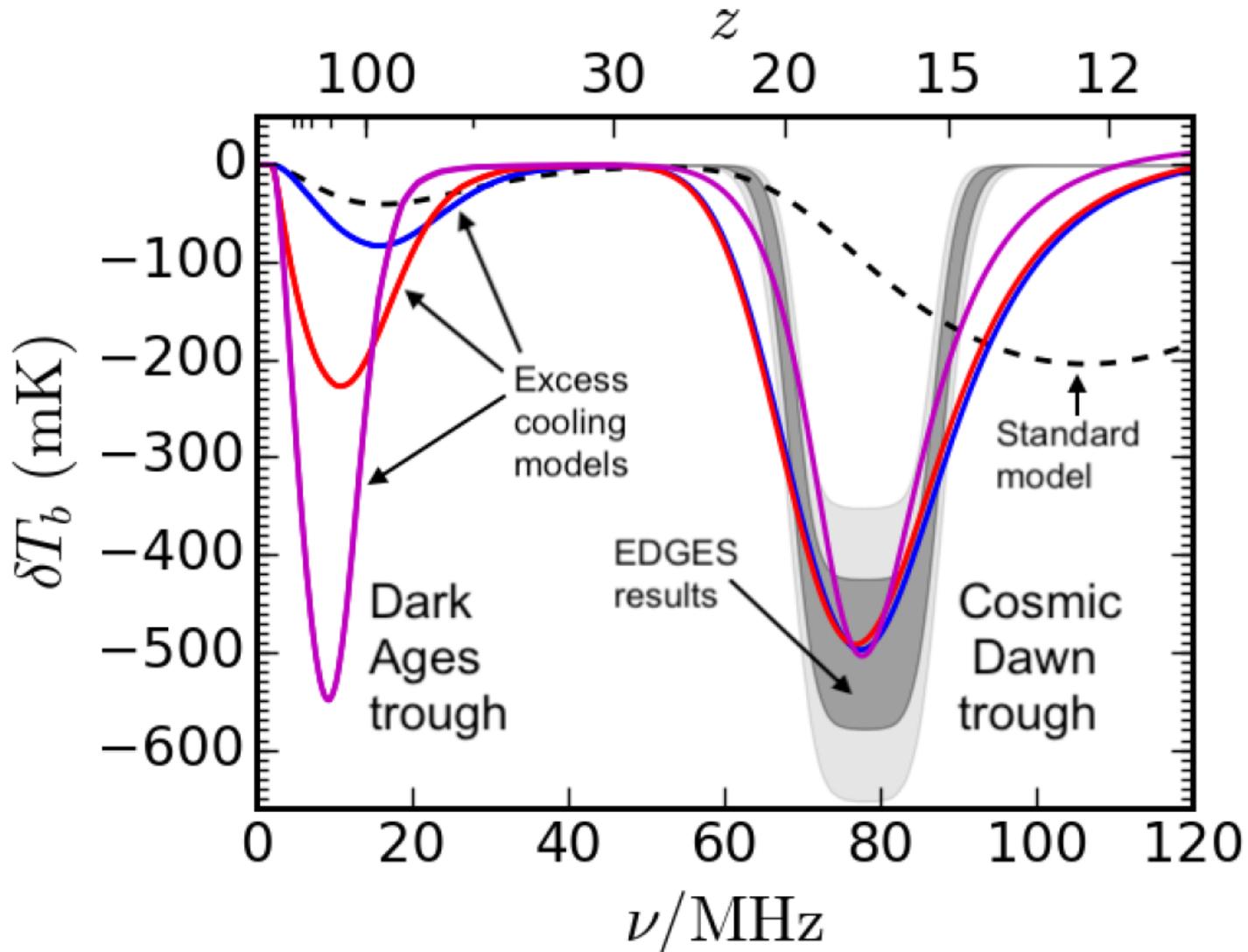


... & D'Angelo 1981

Lamy et al. 2010



The Dark Ages



Requirements

Need many km² of collecting area...

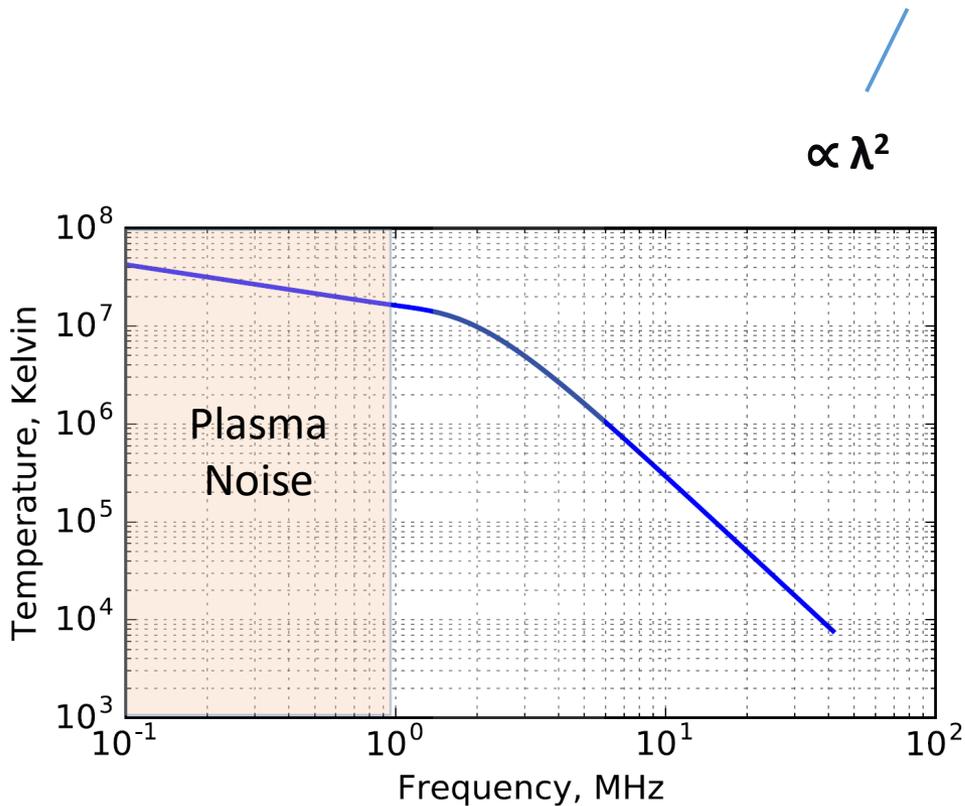
in space...

that can monitor 1000s of stellar systems simultaneously

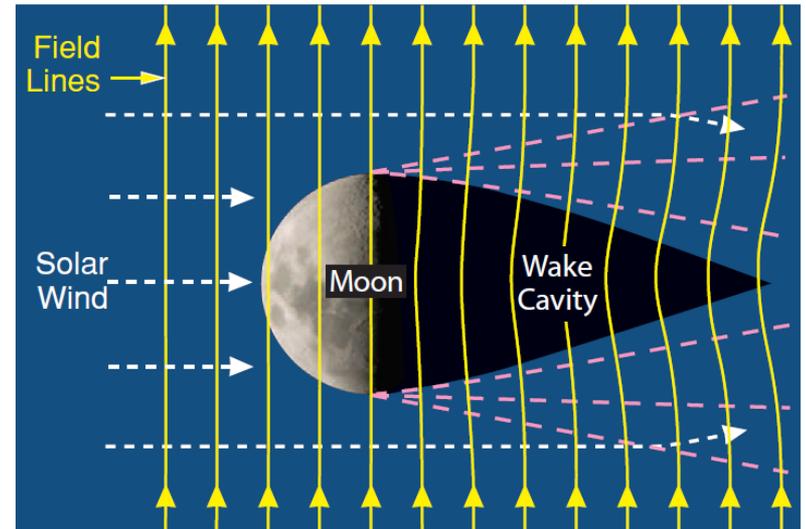
EASY!

The Lunar Farside

Sensitivity of a dipole \propto collecting area / system temperature



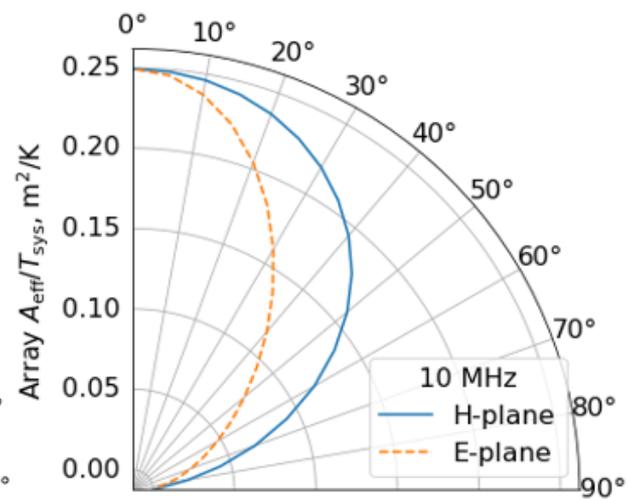
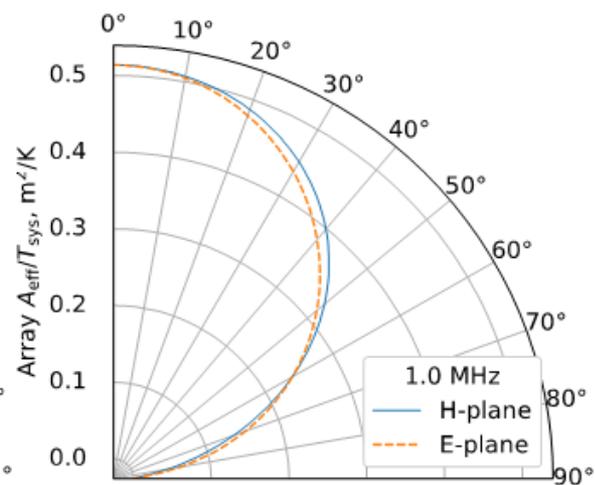
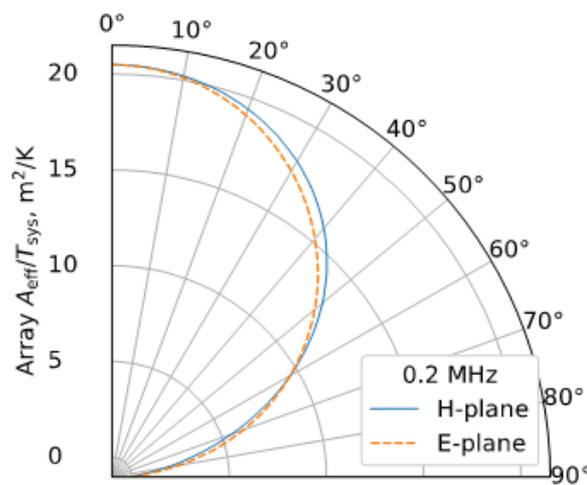
Credit: Andres Romero-Wolf



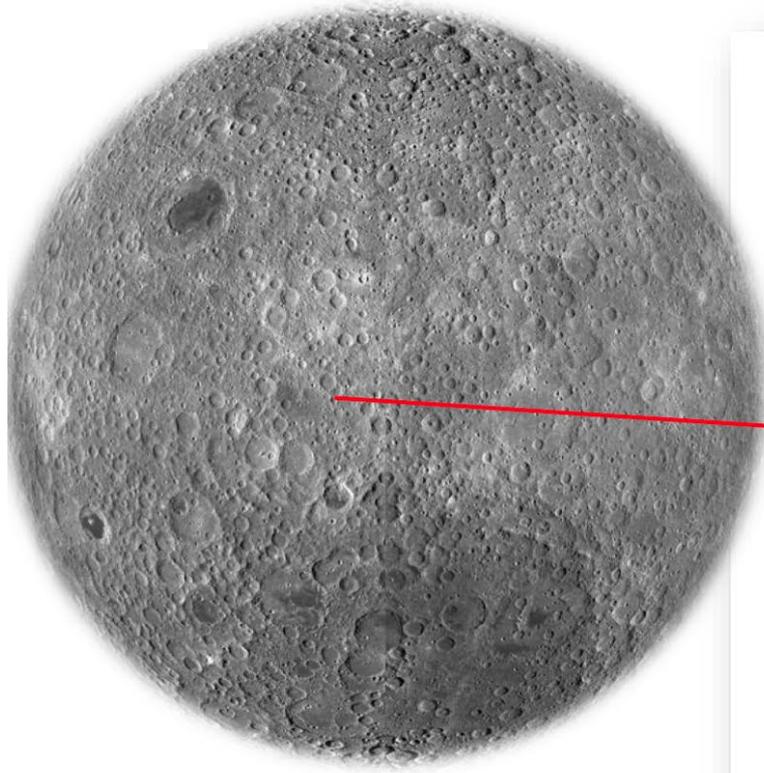
Credit: Steve Bartlett

A dipole of a few meters length on the moon has a collecting area of $\sim 0.3 \text{ km}^2$ at 300 kHz

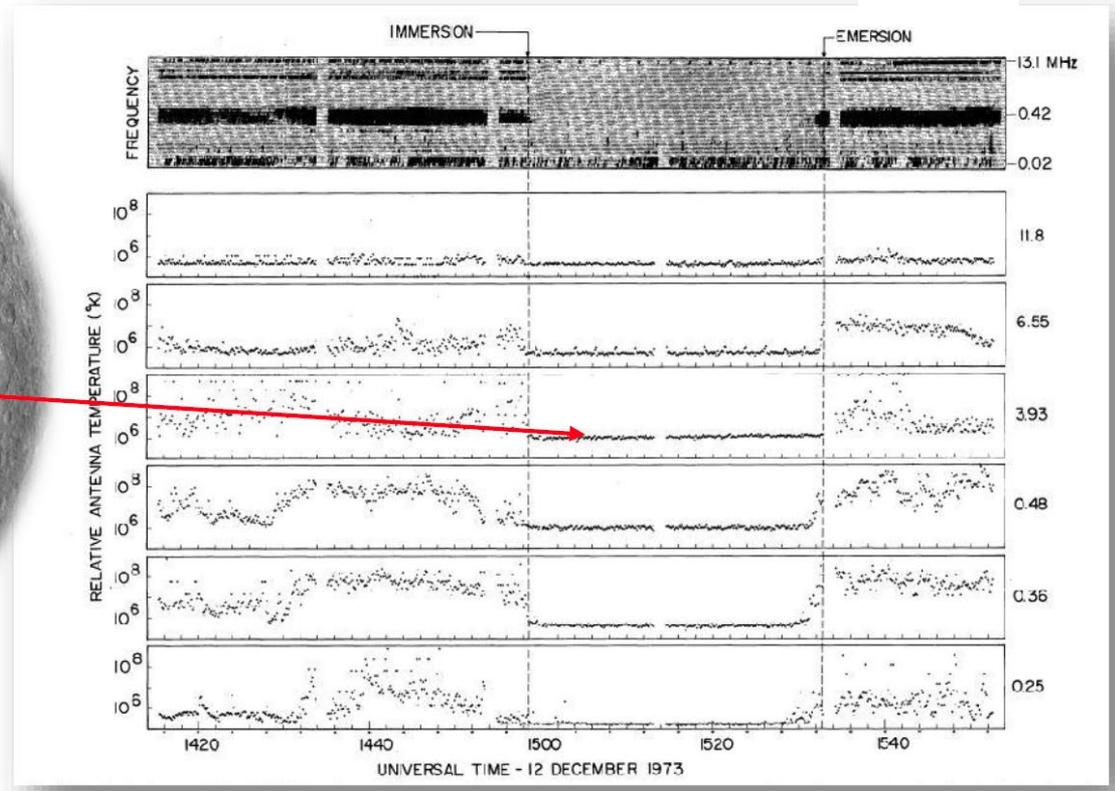
A dipole at 300 kHz is 20x more sensitive than at 30 MHz



Radio-frequency Environment of the Lunar Far-side



RAE-2 1973



RAE-2 occultation of Earth in 1972

FARSIDE Probe Study

- Science Drivers:

The Magnetospheres and Space Environments of Candidate Habitable Exoplanets
The Dark Ages and our Cosmic Dawn

- Assumptions:

- i) Lunar Gateway in operation (available as a communication relay)
- ii) \$1 billion cost cap and 500 kg mass cap [for deployed hardware]

- Timeline:

Nov 2018: Directed probe study commenced

Mar 2019: Overall architecture selected [Team X]

Apr 2019: Follow up mission and instrument studies planned

Jun 2019: Initial report completed

Sep 2019: Engineering Concept Definition Package