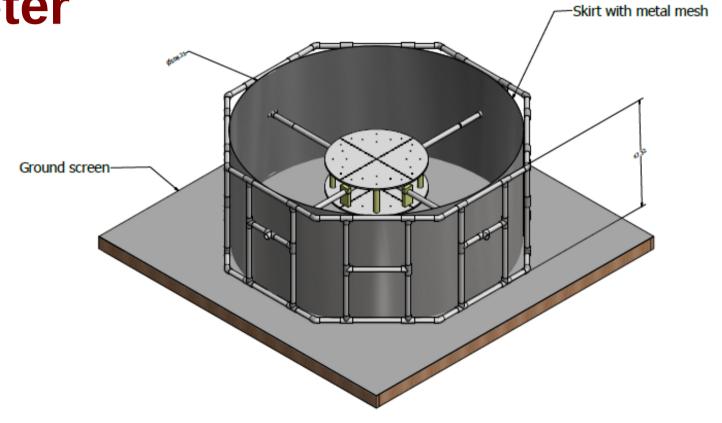


# Cosmic Twilight Polarimeter

A Brief Overview of the Proof-of-Concept System: Instrument, Methodology, and Results





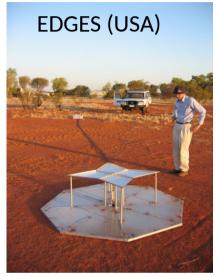
Richard F. Bradley - NRAO CDL & University of Virginia



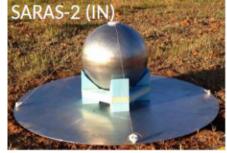
# **Ground-Based 21-cm Global Signal Experiments**

an exciting time for small-scale instruments





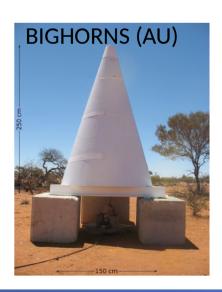
















# **Cosmic Twilight Polarimeter (CTP)**

how does it differ from other global signal experiments?

**Dynamic Polarimetry** 

Tone Injection Based Gain Tracking

Front-End Temperature Stability

Calibration using Models and Lab Measurements

Signal Estimation using Modern Pattern Recognition





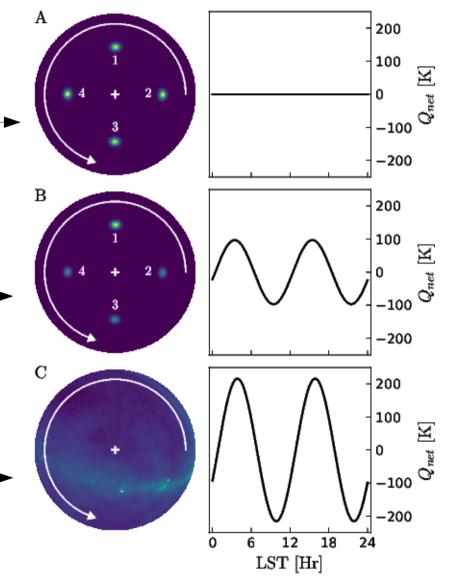
# **Dynamic Polarimetry**

isolating the foreground power spectrum from the 21-cm

Equal flux, equidistant from bore sight

Unequal flux, equidistant from bore sight

Full sky flux, Haslem Map

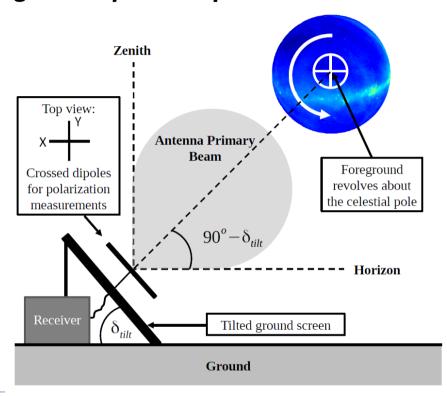






### **Dynamic Polarimetry**

isolating the foreground power spectrum from the 21-cm



#### **Polarimetry Process to measure Foreground**

- 1. Measure "polarization leakage" caused by  $\nu$ -dependence of power patterns of linearly polarized dipoles. Rotation of sky to measure modulated Stokes Q, U, V.
- 2. Harmonic decomposition of modulated Q, U signal, scale to Stokes I, and subtract.

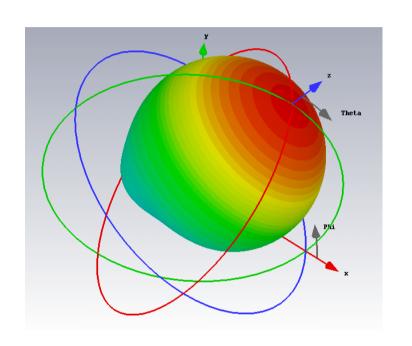


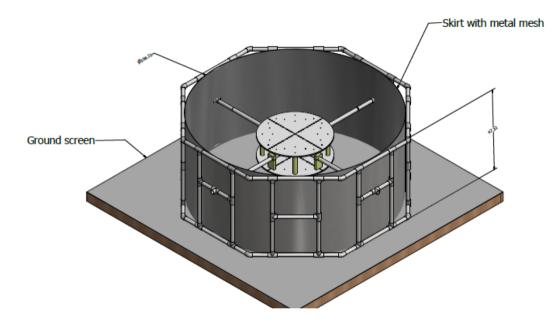
Nhan, Bradley, & Burns, 2017, ApJ, 836, 90.

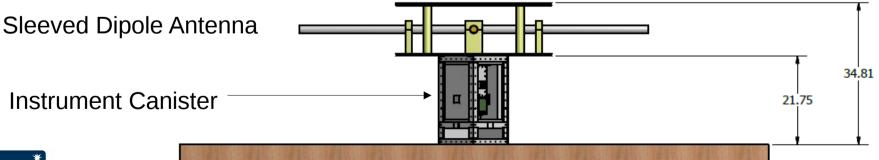


# **Dynamic Polarimetry**

#### with a dual polarization antenna





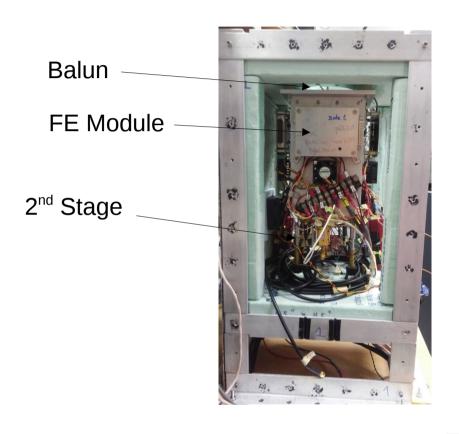


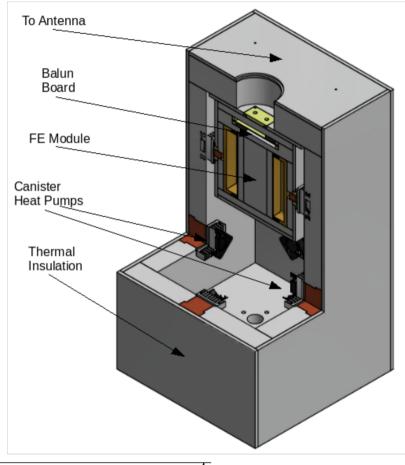


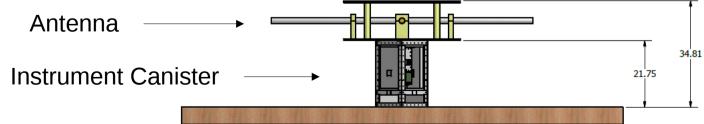


**Radiometer Stability – Temperature Control** 

using two-stage Peltier heat pumps





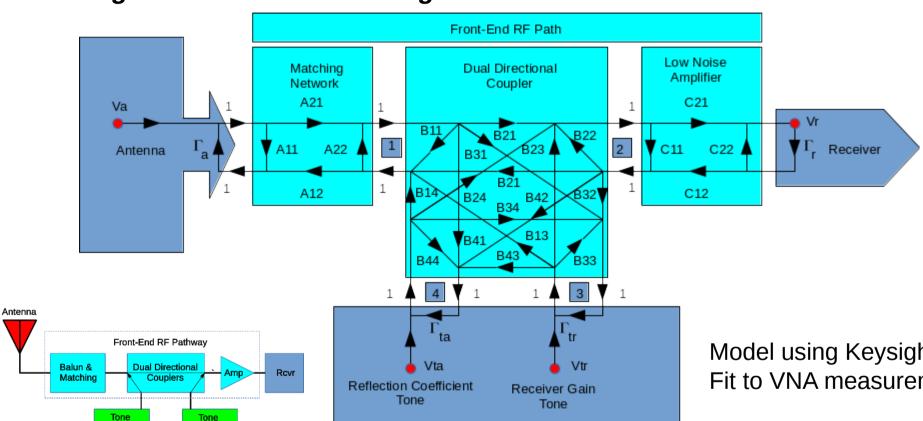






#### Radiometric Calibration: Transducer Gain

through S-Parameter modeling and measurements







Reflection

(Gain)

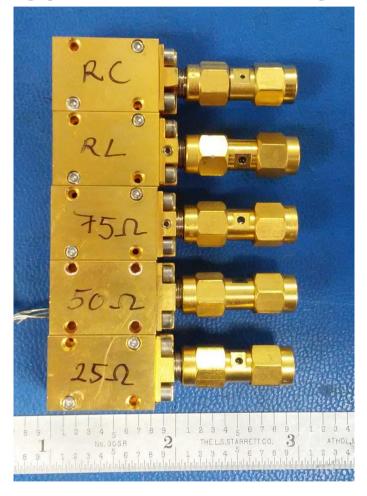
$$G_T = \frac{P_L}{P_{AVS}} = \frac{|S_{21}|^2 (1 - |\Gamma_L|^2)(1 - |\Gamma_S|^2)}{|1 - S_{22}\Gamma_L|^2 |1 - \Gamma_S\Gamma_{in}|^2}$$

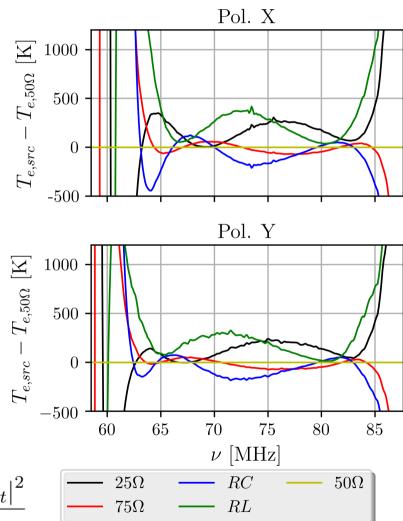
Tone Injection System



## **Radiometric Calibration: Noise Temperature**

using perturbation modeling to estimate noise parameters







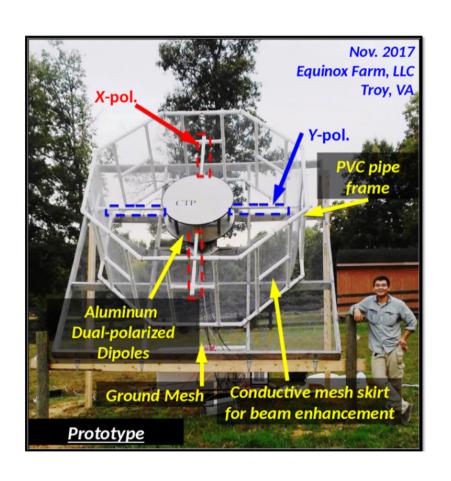
$$T_n = T_{MIN} + \frac{NT_{REF} \left| Z_s - Z_{opt} \right|^2}{R_s R_{opt}}$$

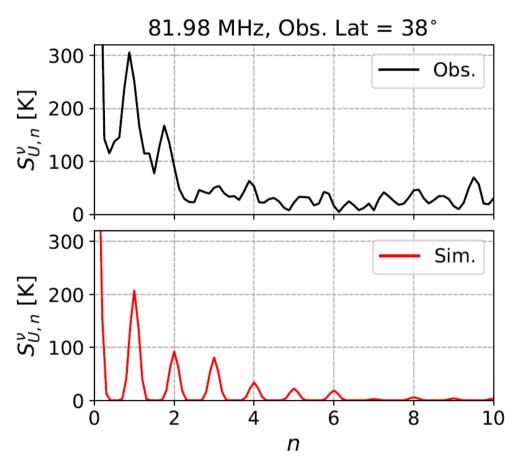
$$\begin{bmatrix} ---- 25\Omega & ---- RC & ---- 50\Omega \\ ---- 75\Omega & ---- RL \end{bmatrix}$$



# Operations in Charlottesville, VA

#### proof of concept and results









# Planned Operations in Green Bank, WV gaining control of systematics

#### CTP deployment with minor upgrades

Bordenave to gain experience with the CTP RFI mitigation activities Reduced RFI environment – radio quiet zone Initial input for SVD algorithm development

#### CTP scaled pair

Use of differential techniques on a pair of scaled antennas Reduce reliance on antenna modeling

#### CTP on the 45' dish

Reduce ground effects
Easily point the CTP



