Data Product 2: m-mode images

Tihkonov regularized m-mode analysis (Eastwood et al. 2018, 2019) (Based on Shaw et al. 2014, 2015)

Cosmic dawn, survey catalogs, Galactic structure, polarization, slow transients

Data Product 3: Astroparticle air showers



First RF-only detection of cosmic-rays (10 events in 40 hours)

Methodology can be applied to detection of tau neutrinos

Monroe et al. 2019



Stage 3: NSF MRI

NSF MRI funding: \$2.2 million

PI team: Gregg Hallinan, Judd Bowman, Dale Gary, Jonathon Kocz, Andrea Isella, Andres Romero-Wolf

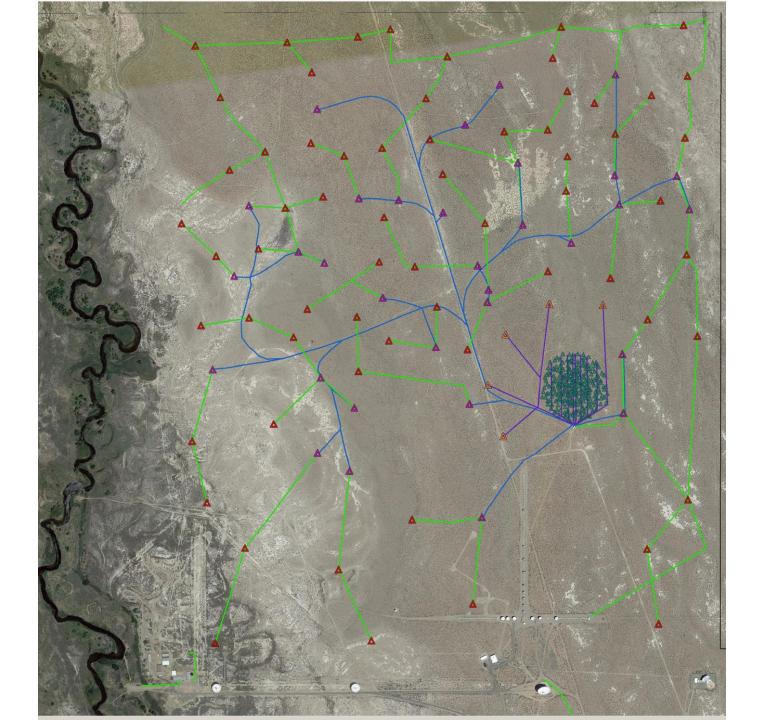
2-year construction effort culminating in a science-ready instrument (88% fabrication)

i) Improved (polarized) imaging performance (better resolution and lower sidelobes)

- ii) Reduced signal path contamination
- iii) Improved calibration performance
- iv) Capability for 1000 hour integrations
- v) Simultaneous all-sky imaging, 12 x beam-forming and cosmic-ray search
- vi) ~5 minute raw voltage buffer
- vii) Integration of 40 m for improved calibration and antenna holography

Baseline: ~100 mJy (1-σ) in 10 seconds at zenith

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Goal:
10 mJy (1-\sigma) in 20 minutes at zenith (sidereal subtraction Stokes I and V)
5 mJy (1-\sigma) in daily m-mode maps (Stokes I and V)
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Mapping Antenna Beams





- Lack of knowledge of the antenna beam is a major limitation for most science and a driver of computational cost (requires peeling)
- Three approaches being implemented in parallel:
 - i) Holography of antennas with 40m dish using pulsar gating
 - ii) beam-mapping via a drone (led by Danny Jacobs of ASU)
 - iii) A novel technique using cosmic-rays

Key Science Projects

Extrasolar space weather Cosmic Dawn Transients Gravitational wave follow-up Solar dynamic imaging spectroscopy Galilean moon subsurface characterization Cosmic rays

