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Functionality-Directed Screening of Pb-Free Hybrid Organic– Inorganic Perovskites with Desired Intrinsic Photovoltaic Functionalities

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(5) Supporting Information

ABSTRACT: The material class of hybrid organic—inorganic perovskites has risen rapidly from a virtually unknown material in photovoltaic applications a short 7 years ago into an ~20% efficient thin-film solar cell material. As promising as this class of materials is, however, there are limitations associated with its poor long-term stability, nonoptimal band gap, presence of environmentally toxic Pb element, etc. We herein apply a functionality-directed theoretical materials selection approach as a filter for initial screening of the compounds that satisfy the desired intrinsic photovoltaic functionalities and might overcome the above limitations. First-principles calculations are employed to systemically study thermodynamic stability and



photovoltaic-related properties of hundreds of candidate hybrid perovskites. We have identified in this materials selection process 14 Ge- and Sn-based materials with potential superior bulk-material-intrinsic photovoltaic performance. A distinct class of compounds containing NH_3COH^+ with the organic molecule derived states intriguingly emerging at band-edges is found. Comparison of various candidate materials offers insights on how composition variation and microscopic structural changes affect key photovoltaic relevant properties in this family of materials.

1. INTRODUCTION

The material class of hybrid organic-inorganic halide perovskites AM^{IV}X^{VII}₃, where A is a small organic molecule, M^{IV} is a low-valent (+2) group-IVA metalloid, and X^{VII} is a halogen, has recently become a rising star in the field of solar-energy conversion materials. With intense research efforts in the past few years,¹⁻¹¹ the power conversion efficiency of thin film photovoltaic (PV) solar cells based on this class of materials has rapidly progressed from the initial value of 3.8% in 2009¹ to a stellar high value of 22.1% only 7 years later.¹² Although most of them are limited to very small cell area ($\sim 0.1-0.2 \text{ cm}^2$),¹³ such efficiencies have become competitive with the record efficiency of the conventional thin film solar cells based on crystalline Si,^{14,15} CdTe,¹⁶ Cu(In,Ga)Se₂,¹⁷ etc. that have been studied for several decades. Quite recently, large area $(>1 \text{ cm}^2)$ solar cells have been fabricated showing conversion efficiency of >15%.¹⁸ Furthermore, hybrid perovskites can be synthesized by inexpensive room temperature solution processing,^{7,9,19-21} making them ideal for low-cost commercial device applications.

Limitations of the Currently Discovered Compounds in Solar Cell Applications. Though exhibiting high power conversion efficiency, the $AM^{IV}X^{VII}_{3}$ family of solar materials is far from optimal, and there exist a number of scientific challenges and practical factors impeding their large-scale device application.^{22,23} These include the following: (i) the devices made by the currently used materials show poor longterm stability under high temperatures and outdoor illumination, as well as in the presence of moisture or oxygen,²⁴ which may be attributed to the intrinsic thermodynamic instability of the materials discovered thus far;^{25,26} (ii) the use of Pbcontaining materials is an environmental concern, and therefore there is a strong desire to replace toxic Pb by a benign element without affecting drastically the conversion efficiency; and (iii) some key properties affecting PV performance need to be further optimized, such as reducing the optical band gap of

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 $[CH_3NH_3]PbI_3$ (1.5 eV)^{27,28} and $[CH_3NH_3]PbBr_3$ (2.35 eV)²⁹ to the optimal value of 1.34 eV from the Shockley–Queisser limit,³⁰ further enhancing light capture efficiency by making the organic molecule contribute in some way to absorption in the solar range, turning material parameters to eliminate anomalous hysteresis in the current–voltage curves,³¹ etc.

Can the Current Limitations Be Overcome by Identifying Other Compounds from the Same General Group? Identification of the alternative $AM^{IV}X^{VII}_{3}$ materials offers a possible solution to overcome the above limitations. Given that this group of compounds is much broader (see Figure 1a) than the most common selection of $A = CH_3NH_3^+$,

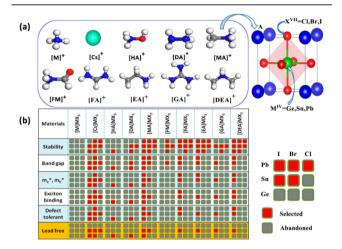


Figure 1. (a) Constitution space of candidate $AM^{IV}X^{VII}_{3}$ perovskites for the materials screening, where 10 cations (see the main text) were chosen for the A site, three group-IVA metalloids (Ge/Sn/Pb) for the M^{IV} site, and three halogen (Cl/Br/I) anions for the X^{VII} site. (b) Step-by-step screening process with the more and more DMs applied (different rows). Each column corresponds to one class of the 9 $AM^{IV}X^{VII}_{3}$ compounds with fixed A, whose arrangement coordinates are shown in the right panel. The red squares mean the materials passing the screening (Selected) and the gray ones mean the materials not passing (Abandoned). The last row of finally winning solar materials is highlighted in orange background.

 $CH(NH_2)_2^+$, $M^{IV} = Pb$ and $X^{VII} = I$, Br, $^{10,11,13,22,24,32-36}$ one also wonders if other members might have advantages and solve the problems encountered by the currently used ones. A few alternative materials have been experimentally synthesized and implemented into solar cells but are showing so far limited conversion efficiency.^{37–39} In this aspect, materials screening via computational simulations is of valuable help to avoid the expensive trial-and-error process of experimental laboratory exploration. To perform the materials screening nonphenomenologically, one needs to establish current understanding of what are the critical, materials-specific "design principles" (DPs) that render such materials superior in solar-energy conversion. To proceed beyond this one needs to formulate the "design *metrics*" (DMs), the embodiment of corresponding DPs, which consist of a series of computable or measurable quantities. In the area of chalcogenide PV materials, successful screening was recently accomplished by applying key properties oriented DMs, computed via first-principles electronic structure theory.^{40,41} In principle, the DMs are not limited to intrinsic materials properties closely relevant to PV performance such as light absorption, defect tolerance, carrier mobility, recombination lifetime, charge extraction, etc. but light capture engineering via larger films thickness or better design of optics (making Si a successful PV platform). However, the application of such methodologies to the hybrid organic—inorganic perovskite family has been historically overlooked and is currently at its early stages.^{42–45}

Here we apply a functionality-directed material discovery approach to screen via systemic first-principles calculations potential high-performance PV hybrid AM^{IV}X^{VII}, perovskites. Such a nonstatistical, science-based inverse-design approach has been recently applied successfully by some of us in theoretical identification and laboratory realization of previously unknown 18-electron Half-Heusler functional compounds.^{46,47} We design the principles/filters that enable selection of the compounds with target photovoltaic performance. The data set of compounds considered for materials screening according to these criteria includes both existing compounds (many not considered yet for photovoltaic applications) as well as new compounds. Particularly, we (i) first formulate the relevant DPs and DMs that would guide us effectively to the compounds with desired optimal PV functionalities and then (ii) search on this basis the promising solar hybrid perovskites satisfying the key compound-intrinsic DMs including thermodynamic stability, as well as the critical properties controlling PV performance such as band gaps, effective masses, dopability, excitons bonding, etc. Figure 1a shows the candidate materials space for screening, whereas Figure 1b provides a summary and conclusion of the screening process where red squares indicate the winning materials that survive the filter of the respective DM. In total, the materials space consists of about 100 compounds, of which as far as we know from literature just 22 were previously reported as being made and 15 have been reported as being considered for PV applications. (There is certainly the possibility that other materials have been experimented with but not reported.) The goal of our materials screening is dual: to identify which materials own potentially superior photovoltaic performance and at the same time to establish the physical-understanding-based principles that can help experimentalists comprehend their both successes and failures. The materials screening process we described led to identification of 14 optimal AM^{IV}X^{VII}, perovskites, as highlighted in the last row of Figure 1b. By comparing various candidate materials, we offer new insights on how composition variation and microscopic structural changes affect key PV relevant properties. Additionally, we find an exotic class of materials with the organic molecule derived state emerging in proximity to band-edges; they are distinct from all the known materials where the molecular state lies far away from bandedges.

We are aware of the fact that even though only 22 of the 100 compounds studied here have been previously reported in the published literature, it is likely that synthetic materials chemists have screened additional compounds that remain unpublished for a variety of possible reasons, including failures to produce good devices at the current stage. We wish to emphasize that the purpose of our theoretical materials screening based on functionality-directed design principles is not just to inform experimentalists on which chemical formulas of such materials they should attempt synthetically next. Ever since the attractive PV performance of Pb-based hybrid perovskites was first noted, this field has progressed mostly by a *trial-and-error* Edisonian approach of combinatorially testing a large number of possible related compounds derived by substitution from a parent compound. Relative successes or failures were sometimes

briefly discussed *ex post facto*. Our aim is actually also to develop understanding-based insights on the physical factors (i.e., the design metrics listed in Section 2) that control both success and failure of compound members in the hybrid perovskite family, whether they were already tried or new ones. We wish that such understanding-based tools can be used as effective guidelines on how to extend the results presented here to search for other PV material systems.

2. DESIGN PRINCIPLES REQUIRED FOR AN OPTIMUM AM^{IV}X^{VII}₃ PV MATERIAL AND THEIR EMBODIMENT AS QUANTITATIVE DESIGN METRICS

The DMs can be divided into (a) the metrics that characterize the $AM^{IV}X^{VII}_{3}$ material itself ("*compound-intrinsic DMs*"), as well as (b) the "*device-level DMs*" specific to the $AM^{IV}X^{VII}_{3}$ material configured within a solar cell device structure that includes components other than this material [such as the hole-blocking layer (HBL) and the electron-blocking layer (EBL), various interfaces and possible metal contacts, etc.]. An instance of the category (b) is the experimentally observed high efficiency of charge collection at the [CH₃NH₃]PbBr₃ absorber/HBL as well as the $[CH_3NH_3]PbBr_3/EBL$ interfaces, creating a natural p-i*n* device where a large fraction of the band gap energy E_{g} is converted to the open-circuit voltage V_{oc} .⁴⁸ As important as this observation is, it however does not reveal why this special absorber material and no other has this superior charge separation in conjunction with the HBL and EBL components. We believe that it makes sense to use the compound-intrinsic DMs as the first filter for preselecting promising candidates. Afterward, the more practical device-level DMs should be used as subsequent filter, applied, however, only to a far smaller group of materials than those screened by the first filter.

We next discuss the basic understanding of the pertinent physical mechanisms controlling PV performance (DPs) that characterize the $AM^{IV}X^{VII}_{3}$ material itself, and the corresponding computable quantities (DMs) that will be used for screening procedures. These *compound-intrinsic DMs* are divided into those that correspond to the structurally ideal $AM^{IV}X^{VII}_{3}$ compounds [items 1–5 below], and those that relate to defects and other inhomogeneities embedded in the compounds [items 6 and 7].

(1) Thermodynamic Immunity toward the Main Decomposition Channel into Competing Phases. The need for a stable structure is one of the first, most obvious conditions required. In the context of the AM^{IV}X^{VII}₃ perovskite family, there are at least two DPs pertaining to stability issues. First, thermodynamic stability requirement needs the eligible PV materials to be stable against all the possible decomposition pathways at solar cell working conditions. The most important decomposition channel of $AM^{IV}X^{VII}_{3}$ is $AM^{IV}X^{VII}_{3} \rightarrow AX^{VII} +$ $M^{IV}X^{V\bar{1}I}_{2}$. The more stable the material, the larger positive magnitude of the decomposition enthalpy (energy difference between the decomposed products and AM^{IV}X^{VII}₃). We thus will use the computed decomposition enthalpy (Figure 2) as the DM here to screen stable compounds. Note that although the decomposition enthalpy is not a direct indication of the materials stability at high temperature, under illumination or moisture etc. required for solar materials, it in principle relates to such stabilities. Additionally, apart from the thermodynamic stability, whether the perovskite materials are formed can be judged by the criterion of crystallographic stability. It can be expressed by two dimensionless parameters: the Goldschmidt tolerance factor t $(t = (r_A + r_X)/\sqrt{2(r_M + r_X)})$ and the

octahedral factor μ ($\mu = (r_M/r_X)$), where r_A , $r_{A\nu}$ and r_X are ionic radii of corresponding ions. Depending on temperature the AM^{IV}X^{VII}₃ family may be stabilized in a simple cubic perovskite or its distorted pseudocubic structure,⁴⁹ and the *t* and μ describe the stabilities of cubic (or pseudocubic) perovskite framework and octahedral M^{IV}X^{VII}₆ units, respectively. To address this crystallographic stability, we will analyze all the candidate materials with respect to *t* and μ (Figure 3) and delineate whether and how feasibly they can be stabilized in the perovskite structure.

(2) Direct Band Gap Matching Solar Spectrum and Strong Absorption near Threshold Due to the p-pEffect. Given a stable material, the solar cell application requires a high solar-energy absorption efficiency, i.e., an optical band gap (absorption threshold) with the suitable magnitude of 1.0-1.7 eV and strong takeoff (slope of absorption versus energy) near threshold. So far all the reported AM^{IV}X^{VII}, materials show strong optical transition at the optical band gaps.³⁶ The physical mechanism responsible for this condition is attributed to the existence of a valence band p (VB-p) to a conduction band p (CB-p) channel of interband absorption.⁵⁰ For the high-Z element M^{IV} , the $M^{IV}-s$ state moves down in energy (forming lone-pair states), separating from $M^{IV}-p$ state (Mass-Darwin relativistic effect).⁵¹ This makes the +2 low oxidation state of M^{IV} favorable in the $AM^{IV}X^{VII}_{3}$ family. As the result, the conduction bands have $M^{IV}-p$ character and not *s* character, and the valence bands are mainly from the antibonding hybridization between halogen $X^{VII}-p$ and $M^{IV}-p$ s states (see Figure 5). Such a VB-p to CB-p band transition channel has the higher intensity (due to the larger degeneracy of p states than s states) than the VB-p to CB-s transitions characterizing most of conventional solar materials such as GaAs or CuInSe₂. This type of design principle has also been identified in the context of discovery of pnictide-chalcogenide absorbers of the Cu-Sb-S family,41 where the high-Z element Sb causes its outer valence 5s orbital to become localized and deeper in energy, thus leaving the p channel as the main one available for bonding. This p-p effect ensures the AM^{IV}X^{VII}₃ family strong absorption intensity at the threshold, which greatly favors solar-energy capture efficiency. We will analyze the direct/indirect nature and magnitude of band gaps (Figure 4), examine the appearance of the p-p effect (Figure 5) that can induce strong band gap absorptions, and use them as the DMs to select the compounds.

(3) Simultaneously Low Electron and Hole Effective Masses at Band Edges. In solar cell devices, photogenerated carriers should be mobile so that they can be efficiently collected at electrodes. One of main factors determining carrier mobility is the carrier effective mass. Most semiconducting materials have low effective mass of electrons (m_e^*) but relatively high effective mass of holes (m_h^*) . For known AM^{IV}X^{VII}₃ materials, not only m_e^* but also m_h^* show quite low values.⁵² This contributes to ambipolar conductivity and is one of the most important advantages of the AM^{IV}X^{VII}₃ materials showing high PV performance. The basic principle operating here is the effect of antibonding states at VB edges formed by hybridization between halogen $X^{VII}-p$ and M^{IV} lone-pair s states (Figure 5). Such hybridization renders the VB edges broaden and more dispersive, leading to the small m_h^* . This is in contrast to the oxides having a rather narrow VB due to localized O-p states or the transition metal compounds where the narrow VB reflects localized d states, both of which show large m_h^* . This type of physics has been elucidated earlier for binary lead chalcogenides $PbX^{VI,53}$ The explicit feature of the $X^{VII}-p$ and $M^{IV}-s$ hybridization in VB for all candidate materials is analyzed in Figure 5. The results for the computed $DMs-m_e^*$ and m_h^* —are shown in Figure 7. It should be pointed out that we consider here the effective masses as the first filter of the charge-carrier mobility metric. One needs to keep in mind, however, that a more refined mobility metric would also involve the momentum scattering time, which, in the case of polar semiconductors, is most likely governed by Fröhlich interactions. These in turn depend on factors such as the high-frequency value of the dielectric function, which changes with the materials chemical composition.

(4) Sufficiently Low Binding Energy and Large Localization Radius for Electron-Hole Pairs. One of the key factors that determines the solar cell performance is whether the photogenerated electron-hole pairs (i.e., excitons) are effectively separated for collection before their recombination. This is to a large extent determined by the properties of excitons. Very high binding energy (E_B) and too small Bohr radius (α_{ex}) of excitons correspond to strong Coulomb interaction between electron and hole and thus are basically detrimental to the electron-hole separation process. The hybrid AM^{IV}X^{VII}₃ perovskites exhibit rather low E_B and large α_{ex} that are comparable to those of conventional semiconductors. For instance, $E_{\rm R}$ of $[CH_3NH_3]PbI_3$ was reported to depend on temperature and vary in a wide range of 2–50 meV,^{54–58} and α_{ex} is around 30 Å.^{59,60} Such low E_B values, which are comparable to the room temperature thermal energy of 25 meV, are beneficial for a large fraction of photogenerated excitons to dissociate spontaneously into free carriers.¹³ This may be one of predominate facts responsible for the exceptionally long diffusion lengths of carriers observed in this family of materials.^{61,62} It has been established that the excitons in this family belong to the three-dimensional Wannier type.⁶³ We thus compute excitonic properties following the standard Wannier-Mott hydrogenic mode, and use the resulted E_B (Figure 8) and α_{ex} (Figure S1, Supporting Information) as the DM for materials screening.

(5) Defect Tolerant Structures Where Bond Breaking Does Not Produce Deep Levels Due to Antibonding Character. To guarantee high carrier concentration and efficient carrier transport, the qualified solar materials need to be absent of deep defect-derived states that are usually trapping centers for carriers' nonradiative recombination. One design principle to avoid such deep defect levels is exploiting the antibonding characters of CB and VB edges. In most semiconducting materials, band gap is formed at the separation of bonding and antibonding states and thus the CB usually has antibonding character. In the case that the VB also derives from antibonding states, bond breaking associated with formation of both *n*-type and *p*-type defects will produce shallow rather than deep levels in the midgap region. This defect tolerant behavior originating from the antibonding characters of VB and CB has been confirmed in ternary chalcopyrites $(I-III-VI_2)^{64,65}$ as well as in $AM^{IV}X^{VII}_{3.}$. Taking $[CH_3NH_3]PbI_3$ as the example, extensive defect calculations indicate most of the intrinsic point defects form shallow levels.^{36,67,68} To probe this defect tolerant behavior for materials to be screened, we provide a thorough analysis of bonding and antibonding features (especially for VB) in Figure 5.

In addition to isolated point defects noted above, the defect tolerant behavior also occurs in the materials that can sustain unusually stable defect pairs. In the context of chalcopyrites (e.g., CuInSe₂), it was previously discovered that defect tolerance is manifested by the existence of highly stable defect-pairs $[2V_{Cu}^- + In_{Cu}^{2+}]$ that give rise to a series of ordered vacancy structures (e.g., CuIn₃Se₈, CuIn₃Se₅, Cu₂In₄Se₇, etc.).⁶⁴ The analogous highly stable defect pairs $[V_A^- + M_A^+]$ in AM^{IV}X^{VII}₃ are likely, which might exist in high concentrations without producing deep carrier trapping centers, giving rise to series of novel compounds (e.g., AM^{IV}₂X^{VII}₅, AM^{IV}₄X^{VII}₉, A₂M^{IV}₅X^{VII}₁₂, etc.).

(6) Intrinsic Defects Can Lead to Both p-Type and n-Type Doping by Control of Chemical Potentials of Reactants during Growth. Once we established materials stability and absence of spontaneously formed detrimental carrier-trapping centers, we need to explore if the qualified solar materials can be made both *p*-type and *n*-type so as to afford, for instance, a p-i-n structure that is usually required by solar cells. According to recent calculations on [CH₂NH₃]PbI₃^{36,68} vacancy of heavy metal $M^{IV}\left(V_{M}
ight)$ is the predominant intrinsic defect responsible for hole formation and interstitial doping of the molecular entity A (A_i) responsible for electron formation. Depending on the relative supply of required reactants during growth it is possible to pin the Fermi energy near the VBM (where V_M renders the sample *p*-type) or near the CBM (where A_i renders the sample *n*-type), thus accomplishing bipolar conductivity.

(7) Microscopic Reason for Exceptionally Long Carrier Diffusion Length, i.e., Why Electron-Hole Recombination Is Low. The long carrier diffusion length is another important factor making the hybrid AM^{IV}X^{VII}₃ perovskites good solar materials. Though the underlying mechanism remains elusive, it may relate closely to intrinsic properties such as small carrier effective masses, low exciton binding energy, absence of deep carrier trapping center, enhanced Born effective charges and lattice polarization, etc.⁶⁷ Another mechanism slowing electron-hole recombination may also exist. For chalcopyrites, it was discovered that the interface between (a) parent compound CuInSe₂ and (b) the Cu-poor ordered vacancy structures places electrons on (a) and holes on (b), thus separating them into different real-space regions.⁶⁹ This effect minimizes the recombination process and enhances carrier transport. The analogous interfacial structure existing in AM^{IV}X^{VII}₃ is likely the interface between parent phase and the highly stable defect-pair derived ordered structures as mentioned.

The strategy of the present work is to compute the DMs we have distilled from our physical understanding of the relevant DPs at work and then use the *compound-intrinsic DMs* as the first-level filter to screen optimal solar $AM^{IV}X^{VII}_{3}$ perovskites. The *device-level DMs* are postponed to future screening. Among the above compound-intrinsic DMs we will use only those that pertain to ideal bulk compounds [items 1–5] rather than the others that pertain to defected or inhomogeneous compounds.

3. ASSOCIATION OF DESIGN METRICS WITH CANDIDATE AM^{IV}X^{VII}₃ MATERIALS

The Constitution Space of Candidate Compounds Considered. Figure 1a shows the constitution space of candidate materials for screening. It is known that most of halide perovskites have temperature-dependent phase diagrams, i.e., while at high temperatures they form cubic perovskites (most likely stabilized by phonon entropy), and at low temperatures distorted-perovskite or low-dimensional nonperovskite phases occur. Taking CsPbI₃ as an example, the

high-temperature cubic perovskite phase transforms to a yellow, large-gap, nonperovskite δ -phase at temperatures below 315 °C; fortunately, its cubic perovskite phase can be stabilized to room temperature by nanostructuring⁷⁰ or composition engineering (e.g., via partially replacing I with Br, see Beal et al.⁷¹). Experimentally all the high-performance solar cells (based on e.g., [MA]PbI₃, [FA]PbI₃, CsPbI₃) are made from the cubic (or slightly distorted) perovskite phases that are stabilized at room temperature against competing nonperovskite phases. Considering this fact, screening of optimal AM^{IV}X^{VII}₃ in th cubic perovskite structure is certainly valuable, even though some compounds form nonperovskite structures at normal synthesis conditions. Therefore, the cubic perovskite phase is uniformly adopted for all the AM^{IV}X^{VII}, to be screened even though some materials may exist with nonperovskite structure at normal conditions. In total 10 inorganic/organic cations were considered on the A site, 3 group-IVA metalloids (Ge, Sn and Pb) on the M^{IV} site, and 3 halogens (Cl, Br, and I) on the X^{VII} site. The chosen monovalent A-site cations have generally comparable ionic sizes, including Cs⁺, ammonium $NH_4^+(M)$, hydroxylamine NH_3OH^+ (HA), diamine $NH_2NH_3^+$ (DA), methylammonium CH₃NH₃⁺ (MA), formamid NH₃COH⁺ (FM), fromamidinium CH(NH₂)₂⁺ (FA), ethylamine CH₃CH₂NH₃⁺ (EA), dimethylamine NH₂(CH₃)₂⁺ (DEA), and guanidine amine $C(NH_2)_3^+$ (GA). In addition to these, we have also considered the cases with the pseudohalogen anions, i.e., $[BF_4]^-$ and $[SCN]^-\!\!\!,$ occupying the X^{VII} site. In total, the materials space consists of about 100 compounds. To our best knowledge, among these candidate materials only 22 were previously made (including CsPbCl₃,⁷² CsPbBr₃,⁷⁵ CsSnCl₃,⁷⁴ CsGeBr₃,⁷⁵ CsGeCl₃,⁷⁵ [MA]GeCl₃,⁴⁹ [MA]-SnCl₃,⁴⁹ CsPbI₃,⁷⁶ CsSnBr₃,⁷⁷ CsSnI₃,⁷⁸ CsGeI₃,^{79,80} [MA]-GeI₃,^{79,80} [MA]SnBr₃,³⁷ [MA]SnI₃,^{37,39} [MA]PbBr₃,¹ [MA]-PbI₃,¹ [FA]GeI₃,^{79,80} [FA]SnI₃,⁸¹ [FA]PbBr₃,³³ [FA]PbI₃,^{11,33} [FA]PbI 38 [CA]CalCal⁸⁰ and the latter [EA]PbI₃,³⁸ [GA]GeI₃⁸⁰), and the last 15 were considered for PV applications.

The explicit screening process in terms of the above *compound-intrinsic* DMs is summarized in Figure 1b, which clearly shows how the inferior materials were knocked out by considering more and more DMs step by step, and which $AM^{IV}X^{VII}_{3}$ live through multiple screenings as the winning materials. The explicit calculated data of corresponding DMs are listed in Table S1–S6 of Supporting Information.

DM (1): Good Thermodynamic Stability and Crystallographic Stability. Figure 2 shows the DM reflecting thermodynamic stability, the decomposition enthalpy ΔH of AM^{IV}X^{VII}, perovskites [except for AGeCl₃, see Experimental Section (ii)]. The order of small molecules in the x-axis is sorted in terms of their steric size (i.e., r_A). Generally one observes a rather large span of ΔH (~1.0 eV). This indicates the thermodynamic stability of this class of compounds depends strongly on the (A, M^{IV}, X^{VII}) combination. While for Pb (Figure 2a) and Sn (Figure 2b) based compounds the change of ΔH from chlorides to bromides to iodides is small, Ge based bromides exhibit the much stronger stability than that of iodides (Figure 2c). Walking through different cations at the A site, we find that four cations, i.e., Cs⁺, MA⁺, EA⁺, and DEA⁺, show evidently strong stability. The perovskites containing the former two cations have been extensively investigated, but the studies of the latter two are rarely reported. 28,38,49 The common feature of the three organic molecules is the appearance of CH₃-terminal. For the prototype material [MA]PbI₃, our calculated ΔH is -0.02 eV/f.u., indicating a

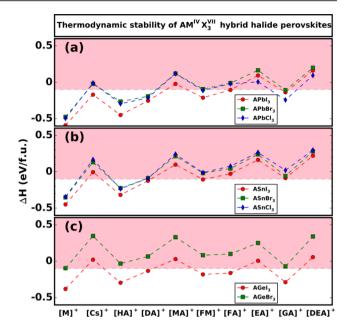


Figure 2. Calculated decomposition enthalpies ΔH of (a) Pb, (b) Sn, and (c) Ge based $AM^{IV}X^{VII}_{3}$ perovskites with respect to decomposed products of $AM^{IV} + M^{IV}X^{VII}_{2}$. Positive ΔH values mean no decomposition occurring. The compounds located in shaded area (with $\Delta H > -0.1$ eV/f.u.) pass the materials screening.

marginal instability against decomposition. The result is qualitatively in accord with the calculation of Zhang et al. where the even lower ΔH of -0.12 eV/f.u. was reported.²⁵

The quantitive discrepancy may originate from the difference in specific structures of decomposed products, van der Waals functionals, orientation of molecules, etc. adopted in the two calculations. Both of the results show consistency with the observed intrinsic thermal instability in experiment, i.e., significant decomposition occurring in [MA]PbI₃ at ~80 °C.²⁶ For all the $AM^{IV}X^{VII}_{3}$ materials, the magnitudes of ΔH (below 0.5 eV/f.u.) are in general much smaller than those of oxide perovskites (up to $\sim 1.5 \text{ eV/f.u.}$).^{82,83} By considering the boundary thermodynamic stability feature of this family, we used a not strict criterion, i.e., $\Delta H > -0.10 \text{ eV/f.u.}$, for the materials screening. This filter makes us discard all the compounds containing the smallest molecule NH₄⁺ and most of ones with HA⁺. 50 compounds with satisfied thermodynamic stability pass the screening to the next step (as in the second row of Figure 1b).

A few notes on the meaning of stability/instability with respect to the disproportionation need to be made: (i) The compounds that are predicted to be slightly unstable with respect to the disproportionation into $AM^{IV} + M^{IV}X^{VII}_{2}$, e.g., the simplest compound of CsPbI₃, might still be synthesized from different reactants other than the above binaries. However, once such a ternary AM^{IV}X^{VII}₃ compound is prepared, it will be prone to-according to our first-principle calculations-decompose into the binary components noted. (ii) The instability with respect to the $AM^{IV} + M^{IV}X^{VII}$ disproportionation is not the only instability channel possible. Other channels will have to be considered for specific cases. For example, the compounds containing multivalent elements such as Sn (that can exist in either valence 4+ or 2+) might be prone for valence instability-transformation of the ASnX^{VII}, with bivalent Sn to the structures affording tetravalent Sn, i.e., destabilization due to oxidation of metallic Sn from 2+ to

4+.^{84,85} (iii) All the calculations here are performed using the standard cubic perovskite structure, and we thus consider the disproportionation stability of this phase. Note, however, that other disproportionation channels involving noncubic phases can be possible in specific cases such as Ge-based iodide perovskites that have been experimentally synthesized recently,^{79,80} exhibiting, however, substantial distortion from the cubic phase, or the [FA]PbI₃, where the black cubic perovskite phase is prone to transform to a yellow phase of different crystallography.⁸⁶

We then map all the candidate materials onto a twodimensional plot with the tolerance factor t and the octahedral factor μ as variables (Figure 3) for analyzing crystallographic

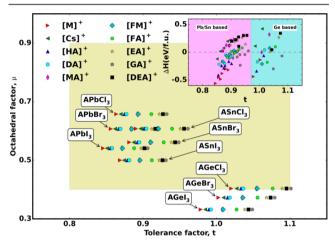


Figure 3. Mapping of various $AM^{IV}X^{VII}_{3}$ perovskites with respect to the octahedral factor (μ) and the tolerance factor (t). Shaded area (with 0.44 < μ < 0.90 and 0.81 < t < 1.11) represents the crystallographic stability requirement for the formability of halide perovskites.⁸⁵ The inset shows the decomposition enthalpies ΔH of the $AM^{IV}X^{VII}_{3}$ perovskites as a function of t. The regions of Pb/Sn and Ge based compounds are shaded by pink and sky blue, respectively.

stability. For inorganic $AM^{IV}\!X^{VII}{}_3$ halides, a previous statistic analysis indicated the formability of perovskites requires 0.81 < t < 1.11 and $0.44 < \mu < 0.90$.⁸⁷ We find that almost all the materials fall in the stable range required by t, except for [EA, GA, DEA]GeCl₃ that are close to the upper boundary of 1.11. The three compounds have actually been abandoned because of their low ΔH . The inset of Figure 3 shows the ΔH of the AM^{IV}X^{VII}₃ perovskites as a function of *t*. Generally the results can be divided into two groups: the Pb/Sn based materials with t < 0.97 (shaded in pink) and the Ge based ones with t > 0.97(shaded in sky blue). For the former group, we observe a general trend of the larger t, the stronger thermodynamic stability (i.e., the higher ΔH). This resembles the behavior observed in oxide perovskites⁸⁸ and suggests that the larger t below 1.0 is more favorable for stabilizing $AM^{IV}X^{VII}_{3}$ in the perovskite structure. At the rather smaller t (<0.89), for instance for [CH₃NH₃]PbI₃, the distorted pseudocubic perovskites with tilted $M^{IV}X^{VII}_{6}$ octahedrons (called orthorhombic γ or tetragonal β phases) become energetically more stable than the standard cubic perovskite (called α phase) at room or low temperatures.⁴⁹ If the energy difference between the γ/β phases and the α phase is substantial, this extra energy gain from structural distortion may stabilize the material despite the low ΔH of the α phase. This explains many known roomtemperature phases of AM^{IV}X^{VII}₃ existing as the less symmetric

structures.^{28,38,49} Turning to the μ value, while Pb-based and Sn-based compounds are safely located within the stable range, the too small μ of Ge-based compounds implies that they might not tend to be stabilized in the perovskite form. This seems, however, not consistent with the recent experiments,^{79,80} where Ge-based iodide perovskites have been successfully synthesized, though accompanied by substantial distortion from the cubic α phase. We thus did not rule out as yet the Ge-based compounds at this step.

DM (2): Suitable Direct Band Gap and Strong Absorption near Threshold. The problem of obtaining the correct band gap in such complex organic-inorganic hybrid structures is a difficult one, and we approach this specific point at this time by distilling information from a few compounds for which the band gap is known and determining the adjustable parameters (i.e., the percentage of exact Fock exchange component) for the employed hybrid functional calculations including the spin-orbit coupling effect. We then use the same parameter values for similar compounds to arrive at the band gap values for the new compounds (see Experimental Section). We found most of the AM^{IV}X^{VII}3 perovskites considered show direct band gap (E_a^d) at the R point of the cubic Brillouin zone. Only 16 materials show indirect gaps formed by the VBM at the R point and the CBM at the point slightly deviating from R (see Figure S2, Supporting Information). This indirect-gap feature has been proposed to facilitate enhanced carrier radiative lifetime.^{89,90} Such indirect gaps are merely several or several tens meV lower than the E_{g}^{d} thus having negligible effect on optical absorption. We thus regard all of them as the directgap materials in the screening process.

Figure 4 shows calculated E_g^d of the AM^{IV}X^{VII}₃ perovskites, and Figure 5 depicts how the band gap is formed by various bonding/antibonding states that are characterized with the analysis of crystal orbital overlap population (COOP).⁹¹ There exists a wide distribution of gap values ranging from about 1.0 to 6.0 eV. This wide variation of E_g^d is attributed to (i) the energy difference between the X^{VII}-p orbital and the M^{IV}-p/s

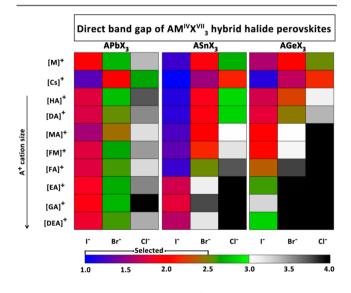


Figure 4. Calculated direct band gaps E_g^d of the AM^{IV}X^{VII}₃ perovskites. The A⁺ cations are sorted by the increasing steric sizes. The gap values of 1.0–3.0 eV (promising for solar materials) are depicted by the RGB colors, and the oversized ones above 3.0 eV are shown in the white-black scale. The criterion of $E_g^d < 2.5$ eV is used for the materials screening.

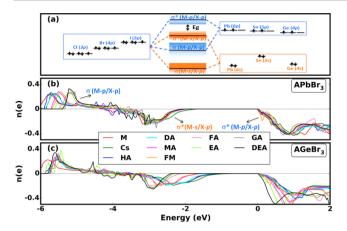


Figure 5. (a) Schematic diagram of constituted atomic orbitals and their formed bonding (σ) /antibonding (σ^*) bands in the AM^{IVXVII}₃ perovskites. (b, c) The crystal orbital overlap populations (COOP) analysis for APbBr₃ and AGeBr₃, respectively. For comparison, the CBM of each material is set to energy zero.

orbitals forming band-edge changes with different ($M^{\text{IV}}\!,\,X^{\text{VII}}\!)$ combinations and (ii) the absolute positions of the CBM and the VBM, which are both antibonding states, being tuned by varied M^{IV}-X^{VII} bonding strengths (see Figure 5a). With increasing electronegativity of X^{VII}, the band gaps show significant enlargement, which is consistent with experimental observations.^{27–29,48,92–94} This makes the majority of the Pbbased and part of the Sn/Ge-based bromides and chlorides not favorable as solar materials owing to their too large E_{σ}^{d} . The iodides family thus contains the most numerous optimal solar materials. Walking through different organic molecules, we found Pb-based compounds show weak dependence of E_{σ}^{d} on A⁺ cations, whereas Sn/Ge-based ones have dramatically increased E_g^d with increasing steric sizes of A⁺. Taking the bromides as the instance, the E_g^d of Pb-based compounds varies within a relatively narrow energy range of 1.98-2.74 eV, but the E_{α}^{d} of Sn and Ge-based ones show the much wider tunability within the range of 1.53-3.77 eV and 1.64-4.15 eV, respectively. Since A⁺ cations have no direct contribution to band-edge states, their influence on E_g^d is exerted though geometric modification of the MIV-XVII network composed of corner-sharing M^{IV}X^{VII}₆ octahedra.

To unravel how the A⁺ cations affect the band gaps of AM^{IV}XV^{II}₃ perovskites, we perform further analysis on the obtained structure-property (i.e., E_g^d) data. We find a geometric quantity of the $M^{IV}X^{VII}_{6}$ octahedron—the averaged bond length of six $M^{IV}-X^{VII}$ bonds (d_{ave})—is a good descriptor for establishing the anticipated structure-property relationship. Figure 6 shows the dependence of E_g^d on d_{ave} for all the $AM^{IV}X^{VII}_3$ materials. One clearly observes that for each (M^{IV}, X^{VII}) combination the magnitude of E_g^d is linearly proportional to the d_{ave} , which actually increases with the steric size of A⁺. The trend becomes more evident from Pb, Sn to Ge-based materials. The physical mechanism responsible for this trend is revealed by projecting band-edge states onto the bonding/ antibonding orbitals in Figure 5. With the increasing sizes of the A⁺ cations of, for instance, Ge based compounds (Figure 5c), the resulted lattice expansion makes the bond lengths of M^{IV}- $X^{\rm VII}$ bonds increase, thus reducing the bond strengths. This pulls down the energy levels of both of the two antibonding states, $\sigma^*(M^{IV} - s/X^{VII} - p)$ (forming VB) and $\sigma^*(M^{IV} - p/X^{VII} - p)$ $X^{VII} - p$) (forming CB). Owing to the more delocalized nature

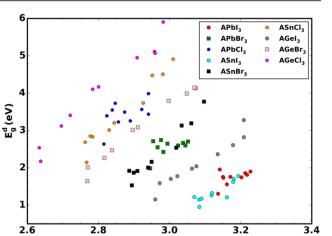


Figure 6. Dependence of direct band gaps E_g^d on the averaged $M^{IV} - X^{VII}$ bond length (d_{ave}) of the $AM^{IV}X^{VII}_{3}$ perovskites.

d_{ave} (Å)

of the $M^{IV} - s$ state than that of the $M^{IV} - p$ state, the interaction between the $M^{IV} - s$ and $X^{VII} - p$ states is stronger, evidenced by the larger energy splitting between $\sigma(M^{IV} - s/X^{VII} - p)$ and $\sigma^*(M^{IV} - s/X^{VII} - p)$ (about 7 eV) than that between $\sigma(M^{IV} - p/X^{VII} - p)$ and $\sigma^*(M^{IV} - p/X^{VII} - p)$ (about 5 eV). As the result the $\sigma^*(M^{IV} - s/X^{VII} - p)$ state is affected more pronouncedly, which exhibits the larger downshift. This enlarges E_g^d and gives rise to its nearly linear increase with d_{ave} . For the Pb-based compounds (Figure 5b), since the $M^{IV}X^{VII}_{6}$ octahedron is relatively compact with the stronger stiffness owing to the larger size of Pb²⁺, d_{ave} is less affected by the changes of A⁺ cations. Thus, E_g^d shows an inconspicuous increase.

From Figure 5 (and Figure S3 in Supporting Information) we can see that all the AM^{IV}X^{VII}₃ perovskites have the VB from the $\sigma^*(M^{IV} - s/X^{VII} - p)$ states and the CB from the $\sigma^*(M^{IV} - s)$ $p/X^{VII} - p$) states. There always exist the VB-p to CB-p channels of band gap transition as mentioned in favor of strong optical transition intensity (see also Figure 9). The magnitude of E_{σ}^{d} can thus act as the valid DM to select the optimal solar materials with high light absorption efficiency. Considering the visible light region of ~1.5–3.0 eV, the criterion of E_a^d lower than 2.5 eV is used for the materials screening. In total 41 compounds pass the screening (Table S3, Supporting Information). This leads us to further select 22 compounds from the 50 winning compounds surviving from the DM (1)(as in the third row of Figure 1b). Among them six are Pbbased compounds, five are Ge-based, and the remaining ones are Sn-based.

DM (3): Low Electron and Hole Effective Masses Favoring Ambipolar Conductivity. As demonstrated in Figure 5, the VB of the $AM^{IV}X^{VII}_{3}$ perovskites is composed of the antibonding states hybridized by the $M^{IV} - s$ and $X^{VII} - p$ orbitals ($\sigma^*(M^{IV} - s/X^{VII} - p)$). This hybridization causes a rather dispersive VB, and thus low hole effective mass (m_h^*). As in most semiconducting materials, the CB dominated by cationic $M^{IV} - p$ states usually gives a small electron effective mass (m_e^*). The feature of simultaneously low m_h^* and m_e^* for this family of materials has been indicated in Figure 7. Most compounds except for [FM] $M^{IV}Cl_3$ ($M^{IV} = Ge,Sn,Pb$) (to be discussed below) show comparably small m_h^* and m_e^* at the order of several tenths of m_0 . This is beneficial for the ambipolar conductivity favored by PV applications. Generally,

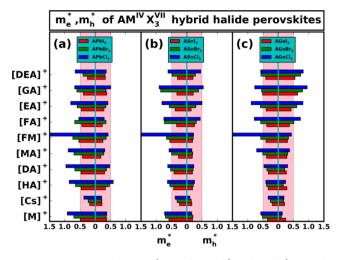


Figure 7. Calculated electron $(m_e^*, \text{ left panels})$ and hole $(m_h^*, \text{ right panels})$ effective masses of (a) Pb, (b) Sn, and (c) Ge based AM^{IV}X^{VII}₃ perovskites. Shaded areas indicate the screening criterion applied $(m_e^* < 0.5 m_0 \text{ and } m_h^* < 0.5 m_0)$.

the values of m_e^* are larger than those of m_h^* , which is distinct from the usual semiconductors with the lighter m_e^* but the heavier m_h^* . Both m_h^* and m_e^* show increasing trends from iodides to bromides to chlorides. For Sn and Ge based compounds, we see a clear trend of m_h^* increasing with the sizes of A⁺ cations. These follow the same trend of E_g^d (Figure 6), which can be understood in terms of the nature of bonding/ antibonding states forming band-edges as above. With screening criteria of $m_e^* < 0.5 m_0$ and $m_e^* < 0.5 m_0$ applied, 40 compounds pass the screening (see Table S4, Supporting Information). Among the 22 winners passing the DM (1, 2), [MA]PbBr₃, [HA]SnBr₃, and [DA]SnBr₃ are ruled out and 19 compounds are left (as in the fourth row of Figure 1b).

DM (4): Low Exciton Binding Energy Not Impeding Photon-Induced Charge Separation. The calculated binding energy (E_B) and Bohr radius (α_{ex}) of Wannier-type excitons for the AM^{IV}X^{VII}₃ perovskites are shown in Figure 8 and Figure S1 in Supporting Infomation, respectively.

For [MA]PbI₃, using the calculated reduced exciton mass μ^* (0.13) and high-frequency limit of dielectric constant ε_{∞} (5.7), we obtain E_B and α_{ex} of 52 meV and 24 Å, respectively. With the ε_{∞} used, in principle the excitonic properties are evaluated for the exciton appearing immediately after photon absorption (i.e., in absence of ion relaxation). The value of E_B reasonably falls in the range of the experimental values^{54,63,95,96} and previous calculations.⁵⁸ One sees E_B shows clear increase from iodides to bromides to chlorides. The trend is in accordance with the observation in experiments.⁶³ This originates from the heavier m_h^* and m_e^* (Figure 7, thus the larger μ^*) and the lower ε_{∞} (corresponding to the larger E_g^d in Figure 4) for the perovskites with the more electronegative halogens. For Sn and Ge based compounds, E_B unambiguously increases with the size of the A^+ cation as the result of the increased m_h^* and the decreased ε_{∞} . To select the materials for facilitating photoninduced electron-hole pairs separation, we use the criterion of $E_B < 100$ meV, which corresponds to $\alpha_{ex} > 15$ Å. This results in the abandonment of most of chlorides and 42 compounds pass the screening (see Tables S5 and S6, Supporting Information). Among the superior compounds surviving the above DMs (1 -3), $[EA]PbI_3$ is ruled out, and 18 are kept as the optimal solar materials (as in the fifth row of Figure 1b).

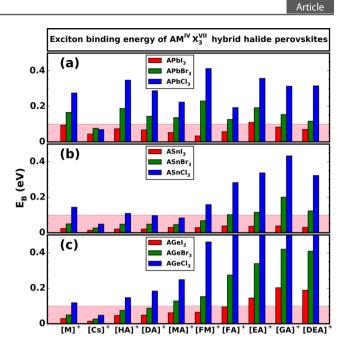


Figure 8. Calculated exciton binding energies of (a) Pb, (b) Sn, and (c) Ge based $AM^{IV}X^{VII}_{3}$ perovskites with the hydrogen-like Wannier– Mott model. Shaded areas indicate the criterion applied ($E_B < 0.1 \text{ eV}$) for the materials screening.

DM (5): Defect Tolerant Behavior Beneficial from the Antibonding Feature of Band-Edge States. As shown in Figure 5 and Figure S3 in Supporting Information, all the studied AM^{IV}X^{VII}₃ perovskites show the VB and CB composed of antibonding states, i.e., $\sigma^*(M^{IV} - s/X^{VII} - p)$ (of VB) and $\sigma^*(M^{IV} - p/X^{VII} - p)$ (of CB). Therefore, they all belong to the materials with the defect tolerant feature as mentioned and pass the screening in terms of this DM. The number of selected superior materials is kept unchanged (as in the sixth row of Figure 1b).

We should mention whereas all the studied $AM^{IV}X^{VII}_{3}$ perovskite compounds show the VB and CB to be composed of antibonding states, so they can all be classified as the defect tolerant materials, expecting shallow defect (donor or acceptor) transition levels, specific calculations on the defect formation enthalpies for particular compounds will be needed in the future. This will determine if some charge-carrier trapping defects can form in large quantities (i.e., have low formation enthalpies) and thus lead to short nonradiative lifetime, which limits photoinduced charge extraction.

In addition to the $AM^{IV}X^{VII}_{3}$ perovskites with halogen anions sitting at the X^{VII} site, we have also investigated the cases involving pseudohalogen anions. Two types of pseudohalogen anions, $[BF_4]^-$ and $[SCN]^-$, are considered. Both of them have been experimentally incorporated into [MA]PbI₃ to partially substitute I⁻ anions.^{97,98} The Pb²⁺/Sn²⁺ and Cs⁺/MA⁺/FA⁺ are chosen as the M^{IV} and A-site cations, respectively. Our calculations (Table S7, Supporting Information) indicate that the materials containing $[BF_4]^-$ show strong stability with respect to decomposition (with $\Delta H > 0.30$ eV/f.u.); the stability of [SCN]⁻ based materials depends on the choice of Asite cations: while the ones with the inorganic Cs^+ are quite stable (with $\Delta H > 0.30$ eV/f.u.), the ones with organic molecular MA⁺/FA⁺ exhibit bad stability (with $\Delta H < -0.15$ eV/f.u.). The $[BF_4]^-$ based materials show extremely large band gaps (even above 8 eV), consistent with previous calculation.⁹⁹ This may be attributed to the high electronegative nature of F that makes the $M^{IV}-X^{VII}$ bond predominantly ionic, weakens the orbital hybridization between M^{IV} and X^{VII} , and thus leads to narrow-width bands and enlarged band gaps. The band gaps of the [SCN]⁻ based materials are not as high as those of the [BF₄]⁻ based ones but still much higher than the upper limit of visible spectrum. Corresponding to the large band gaps, both [BF₄]⁻ and [SCN]⁻ based materials show heavy m_h^* and m_e^* (1.0–8.0 m_0). Therefore, by themselves they are not promising solar materials. By considering the robust thermodynamic stability of [BF₄]⁻ and Cs⁺-[SCN]⁻ based materials, alloying them with other AM^{IV}X^{VII}₃ perovskites may be a promising strategy to improve the stability of currently used materials.

4. EMERGING OPTIMAL SOLAR MATERIALS WITH DESIRED TARGET FUNCTIONALITIES

The above process leads to 18 superior solar $AM^{IV}X^{VII}_{3}$ perovskites that become winners from the materials screening. Of them 9 were previously made but not recognized as PV materials, and 5 were not made. In particular, these materials are chosen in terms of the following criteria: (i) relatively high thermodynamic stability, the decomposition enthalpy being higher than -0.1 eV/f.u., (ii) high light absorption effeciency, the optical band gap sitting below 2.5 eV, (iii) advantages for ambipolar carrier conduction, both electron and hole effective masses being smaller than 0.5 m_0 , (iv) facilitated photon-induced charge dissociation, with small exciton bonding energy (<100 meV), and (v) defect tolerant feature, meaning the appearance of antibonding states in both valence and conduction bands.

Four materials are Pb-based, among which [MA]PbI₃ and [FA]PbI₃ have been experimentally identified as high-performance solar materials with the maximum conversion efficiencies of 19.3%⁸ and 20.1%.¹¹ If we further consider the environmental concern of toxic Pb and apply the screen criterion of Pb-free, they should be excluded from the optimal materials list. Then we obtain 14 optimal materials (as in the last row of Figure 1b), which include 5 Ge-based materials (i.e., CsGeI₃, CsGeBr₃, [HA]GeBr₃, [DA]GeBr₃, and [MA]GeI₃) and 9 Snbased ones (i.e., CsSnI₃, CsSnBr₃, CsSnCl₃, [MA]SnI₃, [MA]SnBr₃, [FA]SnI₃, [EA]SnI₃, [GA]SnI₃, and [DEA]SnI₃). Among them, CsSnBr₃, CsSnCl₃, CsGeBr₃, [MA]SnI₃, [MA]-SnBr₃, [EA]SnI₃, and [DEA]SnI₃ own the much higher ΔH (above 0.1 eV/f.u.) than that of $[MA]PbI_3$ (-0.02 eV/f.u.), indicating their potentially outstanding thermodynamic stability. The attempt to synthesize them and use them for PV applications holds promise for resolving the issue of poor longterm stability in the solar cell devices based on [MA]PbI₃. The band structures and absorption spectra of selected optimal materials are shown in Figure 9. As seen all these materials show similar band structures in the band gap region, e.g., with the direct gap forming by the dispersive VBM and CBM at the R point (Figure 9a). Most of them show comparably strong absorptions at the threshold of between 1.0 and 2.0 eV (Figure 9b), implying potentially high solar-energy capture efficiency. Note that although there is an instability issue associated with the tendency of $\text{Sn}^{2+}/\text{Ge}^{2+}$ oxidized to $\text{Sn}^{4+}/\text{Ge}^{4+}$, it could be prevented during materials synthesis or device fabrication. The solar cells made by [MA]SnI₃ have been reported with an early stage efficiency around 6%.37,39

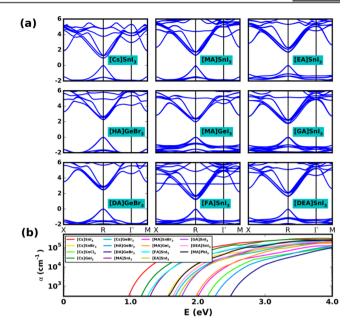


Figure 9. (a) Calculated band structures of selected finally winning $AM^{IV}X^{VII}_{3}$ perovskites passing all the DMs of the materials screening. (b) Absorption spectra of 14 Pb-free winning $AM^{IV}X^{VII}_{3}$ perovskites compared with that of $[MA]PbI_{3}$ (in black).

DISCOVERY OF AN EXOTIC CLASS OF HYBRID AM^{IV}X^{VII}₃ PEROVSKITES WITH BAND-EDGES INVOLVING THE ORGANIC MOLECULE DERIVED STATE

It generally appears that for the known hybrid AM^{IV}X^{VII}₃ materials the role of organic molecular A⁺ cations is stabilizing the perovskite lattice and, thus, has only an indirect effect on the PV relevant properties through modifying the M^{IV}-X^{VII} framework (e.g., the specific orientation of them may cause indirect bandgap and slow carrier recombination⁸⁹). This is because the organic molecule derived electronic states are located far away from band-edges,⁶⁸ thus not participating in visible light absorption, charge carrier separation and transport, etc.

We find here this picture is not always true. Figure 10a shows the band structures of $[FM]GeX^{VII}$, $(X^{VII} = I, Br, Cl)$. Clearly there appears a quite flat band (with red circles) close to the CBM in [FM]GeI₃. With the band gap increasing from iodides to bromides to chlorides, its energy level moves gradually toward the gap region and becomes the real CBM in [FM]GeCl₃. From the projected density of states of [FM]GeI₃ (Figure 10b), we can see it is derived from the FM molecule (red lines). The extremely dispersiveless nature of these localized states is responsible for the abnormally huge m_e^* in $[FM]M^{IV}Cl_3$ ($M^{IV} = Ge, Sn, Pb$) (Figure 7). Note that the relative position of the FM-molecule derived band within the CB relates closely to the relativistic effect of spin-orbit coupling (SOC). While the SOC does not affect much the VB mainly from the M^{IV}-s states, it has significant impact on the CB that is dominated by the M^{IV}-p states. As shown in Figure S4 in Supporting Information, with the SOC included the CB formed by the M^{IV}-X^{VII} framework shows evident downshift, and the shift amount decreases from the heavier Pb, Sn, to the lighter Ge based compounds. Since the FM-molecule derived states are much less affected by the SOC, the energy difference between them and the minimum of the CB formed by the

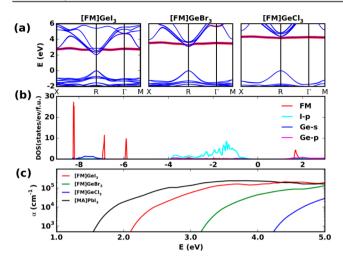


Figure 10. (a) Electronic band structures, (b) projected density of states (DOS), and (c) absorption spectra of $[FM]Ge(I/Br/Cl)_3$. In (a) the orbital projections onto the FM molecule are indicated by red circles. The absorption spectrum of $[MA]PbI_3$ is given in (c) for comparison.

 $M^{IV}-X^{VII}$ framework decreases from Pb, Sn, to Ge based compounds and becomes negative in [FM]GeCl₃ (making the molecular states turn into the actual CBM). The appearance of such molecule derived states in proximity to the CB edge in [FM]M^{IV}X^{VII}₃ is attributed to the relatively low binding energy of the antibonding state of the C=O bond for the FM molecule. These states are strongly localized at the molecule sites. Though in $[FM]M^{IV}X^{VII}_{3}$ the states have negligible effects on the light absorption (Figure 10c) because of the large gap (about 8 eV) between their highest occupied and lowest unoccupied molecular orbitals (Figure 10b) and the forbidden optical transition between them and the states of M^{IV}-X^{VII} framework; finding of such molecule derived states around band edges may represent an advance in having organic molecules involved in light absorption to increase solar capture efficiency. On the other hand, because of the strongly localized characteristic, they may behave as carrier trapping centers mimicking deep defect states and thus impede photoinduced carrier separation and transport to electrodes. This would be detrimental to the solar-to-electric conversion efficiency. Therefore, caution needs to be taken in designing the AM^{IV}X^{VII}₃ perovskites with such states in proximity to band edges.

6. CONCLUSION AND DISCUSSION

In summary, we present via systemic first-principles quantummechanical calculations a *materials-by-design* exploration of high-performance hybrid organic—inorganic perovskites with desired intrinsic photovoltaic functionalities. Based on the knowledge of what makes this family of materials superior in solar cell applications, we formulate the scientific, materialsspecific *design metrics* that consist of a series of computable quantities (e.g., thermodynamic and crystallographic stability, light absorption, carrier effective mass, dopability, exciton binding, etc.) and use them to screen the compounds that satisfy explicitly such target functionalities. We have identified in this materials screening program 18 winning compounds from a materials space composed of ~100 candidates, including the most commonly used materials of $[CH_3NH_3]PbI_3$ and $[CH(NH_2)_2]PbI_3$. Of them 14 are Pb-free (including 5 Gebased materials, i.e., CsGeI₃, CsGeBr₃, [NH₃OH]GeBr₃, [NH₂NH₃]GeBr₃, and [CH₃NH₃]GeI₃, and 9 Sn-based materials, i.e., CsSnI₃, CsSnBr₃, CsSnCl₃, [CH₃NH₃]SnI₃, $[CH_3NH_3]SnBr_3$, $[CH(NH_2)_2]SnI_3$, $[CH_3CH_2NH_3]SnI_3$, [C-(NH₂)₃]SnI₃, and [NH₂(CH₃)₂]SnI₃), and 7 show substantially enhanced thermodynamic stability with respect to [CH₃NH₃]-PbI₃ (i.e., CsSnBr₃, CsSnCl₃, CsGeBr₃, [CH₃NH₃]SnI₃, [CH₃NH₃]SnBr₃, [CH₃CH₂NH₃]SnI₃, and [NH₂(CH₃)₂]- SnI_3). Furthermore, we have discovered a distinct class of hybrid perovskites from the known ones, i.e., the compounds containing NH₃COH⁺, exhibiting the organic molecule derived electronic states in proximity to band-edges. Our results also offer useful insights on the structure-property relationship, i.e., how microscopic structural changes affect photovoltaic relevant properties in this system. The current study provides a roadmap for acceleration of experimental discovery of alternative solarenergy hybrid perovskites to further boost the conversion efficiency and overcome the currently identified limitations of the existing materials. Finally we note that the current work focuses on selection of superior solar hybrid perovskite in the single composition. In reality the excellent photovoltaic performance can be achieved by compositional engineering of multiple perovskite materials, such as the conversion efficiency of 18% reached by mixing [CH(NH₂)₂]PbI₃ with [CH₃NH₃]-PbBr₃.¹⁰ In such cases, our results reported here on thethermodynamic stabilities, optical band gaps, carrier effective masses, etc. of a broad range of hybrid perovskites will essentially provide useful guidance for how to choose suitable materials for the compositional engineering.

Note that experimentally thus far [MA]PbI₃ has been reported to make better performing solar cells than the Snand Ge-based alternatives. Particularly, even though our calculations predict Sn-compounds to be more stable than the equivalent Pb-compounds as far as the tendency to disproportionation reaction $AM^{IV}X^{VII}_{3} \rightarrow AX^{VII} + M^{IV}X^{VII}_{2}$ is concerned (Figure 2), there are credible reports that the Sncompounds often lead to poor-performance solar cells, showing intrinsic *p*-type conductivity and decrease in photocurrent density and fill factor.^{28,37,39} This suggests that additional criteria beyond intrinsic bulk properties may be at play. This is analogous to the replacement in 2CuInS₂ of 2In by Zn + Sn creating Cu₂ZnSnS₄ (CZTS) that involves Sn substitution. Because of the multivalency nature of Sn (being Sn²⁺ as well as Sn⁴⁺), an associated deep defect center detrimental to photovoltaic performance may form.¹⁰⁰ Among AM^{IV}X^{VII}₃ perovskites, while Pb-compounds contain rather stable Pb²⁺, oxidation disproportionation of Sn²⁺ to Sn⁴⁺ is prone to occur in Sn-compunds. The same physics might be at work here and is under investigation.

Whereas indeed some of our selected winning compounds may be known to experimentalists and tried for fabricating solar cells, we must emphasize that the goal of this work is to supplement the "*try-them-all*" Edisonian shotgun approach usually used in this field (as well as in combinatorial chemistry) by an understanding based approach encompassing a list of "principles/filters" needed to make scientific choices and to understand the performance of choices already made. Even if our selected compounds did not work experimentally, our study is still meaningful to the community—the implication is that additional factors beyond the ones involved here need to be considered to reliably evaluate the photovoltaic performance of the AM^{IV}X^{VII}₃ material system. Table I. Calculated Lattice Constants and Band Gaps of $[MA]Pb(I/Br/Cl)_3$ and $[MA](Pb/Sn/Ge)I_3$ at Different Levels of Theory, Compared with the Available Experimental Values^{*a*}

			different levels of theory					
properties	materials	exp.	optB86b	D2	vdw-DF2	optPBE-vdw	DFT	
lattice constant <i>a</i> (Å)	[MA]PbCl ₃	5.67 ¹¹³	5.69	5.67	5.83	5.76	5.81	
	[MA]PbBr ₃	5.90 ¹¹³	5.95	5.91	6.13	6.03	6.07	
	[MA]PbI ₃	6.31 ³²	6.33	6.29	6.54	6.42	6.47	
	[MA]SnI ₃	6.23 ²⁸	6.26	6.24	6.50	6.38	6.44	
	[MA]GeI ₃		6.10	6.08	6.34	6.22	6.33	
			different levels of theory					
properties	materials	exp.	HSE*+SOC	HSE+SOC	HSE	PBE+SOC	PBE	
band gap (eV)	[MA]PbCl ₃	3.11 ²⁹	3.23	2.08	3.18	1.38	2.44	
	[MA]PbBr ₃	2.35^{114}	2.43	1.41	2.50	0.81	1.86	
	[MA]PbI ₃	1.52^{32}	1.55	0.75	1.90	0.27	1.37	
	[MA]SnI ₃	1.20 ²⁸	1.26	0.40	0.75	0.03	0.37	
	[MA]GeI ₃	1.90 ⁸⁰	1.98	1.31	1.41	0.84	0.96	

"For the HSE*+SOC calculations, 55%, 70%, and 55% percentage of exact Fock exchange are used for Pb, Sn and Ge-based perovskites, respectively [see Experimental Section (iii)].

7. EXPERIMENTAL SECTION

All calculations are performed within the framework of density functional theory (DFT) using the plane-wave pseudopotential method as implemented in the VASP code.¹⁰¹ The electron-ion interaction is described by means of projector-augmented wave (PAW) pseudopotentials.¹⁰² Configurations of ns^2np^2 for Pb/Sn/Ge, ns^2np^5 for I/Br/Cl, $2s^22p^2$ for C, $2s^22p^3$ for N, $2s^22p^4$ for O, and 1s for H are considered as valence electrons. The scalar-relativistic effects (including the mass-velocity and Darwin terms), which are especially important to high-Z elements of Pb/Sn, are included. The inclusion of semicore d electrons into the PAW pseudopotentials of Pb/Sn/Ge is found to have negligible effects on the calculated DMs. We use the Perdew-Burke-Ernzerhof (PBE) parametrization of the generalized gradient approximation¹⁰³ as the exchange-correlation functional. A plane-wave basis set with an energy cutoff of 520 eV and Monkhorst-Pack k-point meshes with spacing $2\pi \times 0.05$ Å⁻¹ or less are used for total energy minimization. The convergence of total energy with respect to these parameters is examined and found to reach the level of less than 0.001 eV/f.u. To properly take into account the long-range dispersion interactions that play an important role in the current system involving organic molecules, we have carefully assessed performance of various van der Waals (vdWs) functionals in describing equilibrium lattice parameter of CH₃NH₃-containing materials, i.e., [CH₃NH₃]Pb(I/Br/Cl)₃ and [CH₃NH₃](Pb/Sn/Ge)I₃ in the cubic α phase, as depicted in the upper part of Table I. We find that both optB86b¹⁰⁴ and D2-vdW¹⁰⁵ functionals give satisfied agreements between calculations and experiments (with the maximum deviation below 0.8%). The former is chosen for the energetic calculations and structural optimizations. The hybrid functional (HSE)¹⁰⁶ approach including the spin-orbit coupling (SOC) effect is used for accurately calculating band gaps. Creation of calculation workflows, management of large amounts of calculations, extraction of calculated results, and postprocessing analysis are performed by using an open-source Python framework designed for large-scale highthroughput energetic and property calculations, the Jilin University Materials-design Python Package (Jump2, to be released soon). More specific details on calculating various compound-intrinsic DMs are described as below:

(i) The **structure optimization** of all the AM^{IV}X^{VII}₃ perovskites is done in the α phase with cubic unit cell. We have minimized the total energy with respect to the cubic lattice constant and internal atom positions. The resulted tilting of M^{IV}X^{VII}₆ octahedra from the broken symmetry caused by nonspherical A molecules is taken into account. It is generally accepted that the small A molecule freely rotates in the room-temperature α phase owing to the low rotation barrier.^{89,107,108} We here align the principal axes of small molecules along the $\langle 111 \rangle$ direction that is one of energetically favorable directions.³⁶ The other

aligned direction (e.g., $\langle 100 \rangle$) of small molecules is examined and found to have negligibly minor effect on the compound DMs (see for instance the decomposition enthalpy results in Figure S5, Supporting Information).

(ii) The **decomposition enthalpy** (ΔH) of an $AM^{IV}X^{VII}_{3}$ perovskite is defined as the energy gain during the reaction channel of $AM^{IV}X^{VII}_{3}$ $\rightarrow AX^{VII} + M^{IV}X^{VII}_{2}$. So positive ΔH values indicate no decomposition. This should be the predominant decomposition channel by considering the experimental synthesis routine of $AM^{IV}X^{VII}_{3}$ materials usually involving $M^{IV}X^{VII}_{2}$.^{24,28,38,49} For AX^{VII} , the cesium chloride structure, which is the *e*-phase of [CH₃NH₃]I,^{109,110} is adopted. For $M^{IV}X^{VII}_{2}$, the actual low-temperature structure from experiments is taken; this does not include GeCl₂ that is experimentally absent, and thus the ΔH of AGeCl₃ is not calculated.

(iii) Band gaps: To reliably calculate band gaps of the $AM^{IV}X^{VII}$ perovskites, we examine performance of different levels of theoretical approaches for CH₃NH₃-containing materials, i.e., [CH₃NH₃]Pb(I/ Br/Cl)₃ and [CH₃NH₃](Pb/Sn/Ge)I₃, as summarized in the lower part of Table I. As expected the DFT-PBE approach seriously underestimates the band gaps of all the materials due to the selfinteraction error. The HSE approach (with the standard 25% nonlocal Fock exchange) rectifies to some extent this problem and gives larger gaps than experimental values for all the Pb-based compounds but still underestimates those of the Sn and Ge-based ones. The inclusion of the SOC results in expected significant reduction of gap values for Pb and Sn-based compounds. For all the materials, the HSE+SOC approach gives the lower gap values than the experimental ones. These results are generally consistent with the previous calculations³⁶ and indicate that the higher-level GW calculations with the SOC included are necessary for accurately calculating band gaps of this system.¹¹¹ Unfortunately the SOC-GW method is extremely computationally expensive, which is not affordable for the current materials screening involving hundreds of materials. Taking into account these facts, we use the HSE+SOC method with adjustable percentage (α) of exact Fock exchange (HSE*+SOC). In particular, we adjusted the α values to reproduce the experimental gap values of [CH₃NH₃](Pb/Sn/Ge)I₃ and found that $\alpha = 55\%$ for $[CH_3NH_3]PbI_3$, 70% for $[CH_3NH_3]SnI_3$, and 55% for [CH₃NH₃]GeI₃ give good agreements between theory and experiment (with the maximum deviation of 5%). We then adopt this set of α values for the calculations of other Pb, Sn and Ge-based materials, respectively. This HSE*+SOC approach, e.g., using the unified α for the materials containing the same M^{IV} cation, is validated by the satisfied agreements of gap values between theory and experiment for [CH₃NH₃]PbBr₃ and [CH₃NH₃]PbCl₃.

(iv) Steric sizes of organic cations: To calculate the tolerance factor *t* and the octahedral factor μ , one needs to know the ionic radii of A⁺, M^{IV2+}, and X^{VII-} ions. While the effective crystal ionic radii of

 M^{IV2+} and X^{VII-} (r_M and r_X) are readily known,¹¹² evaluation of the sizes of nonspherical organic A^+ cations (r_A) is challenging. Here we calculate r_A through the expression of within the idealized solid-sphere model, where *a* is cubic lattice constant from our theoretically optimized structures; the ionic radius of inorganic Cs⁺ is used to derive the radii of other organic cations. For instance, the ionic radius of MA⁺ (r_{MA}) is calculated via the relation of $\frac{r_{MA} + r_X}{r_C + r_X} = \frac{a_{(MA)}M^V X_X^{YII}}{a_{CM}N_X_X^{YII}}$, and then the averaged value among all the (M^{IV} , X^{VII}) combinations is taken (see Figure S6, Supporting Information).

(v) The effective mass (m^*) tensors that relate directly to carrier's electrical conductivity are calculated through the semiclassical Boltzmann transport theory.¹¹⁵ The DFT-PBE derived Kohn–Sham eigenvalues at a more dense k-points grid of $2\pi \times 0.016$ Å⁻¹ are used for these calculations. The unified carrier concentration of 1.0×10^{18} cm⁻³ and room temperature of 300 K are chosen. This takes into account effects of nonparabolicity, anisotropy of bands, multiple bands, etc. on carrier transport. For all the AM^{IV}X^{VII}₃ materials in the cubic α phase, the anisotropy of m^* tensors is small, and the averaged m^* among the values of three principal axes is taken as the DMs.

(vi) The exciton bonding energy (E_B) and Bohr radius (α_{ex}) are calculated in terms of the hydrogen-like Wannier–Mott model with the calculated averaged m^* of electron and hole, as well as dielectric constant as the input parameters. Particularly, the E_B is given by $E_B = \mu^* R_y / m_0 \varepsilon_{\infty}^2$, where μ^* is the reduced exciton mass (i.e., $1/\mu^* = 1/m_e + 1/m_h)$, R_y is the atomic Rydberg energy, and ε_r is the relative dielectric constant. Here the high-frequency limit of dielectric constant (ε_{∞}) resulted from electronic polarization, which is calculated from the finite-electric field calculations using the Berry-phase technique, ^{116,117} is adopted as ε_r . The resulted E_B describes the excitons generated immediately after phonon excitation (without involving lattice relaxation). The α_{ex} is given by $\alpha_{ex} = \alpha_H \varepsilon_r \frac{m_0}{\mu^*}$, where α_H is the Bohr radius.

ASSOCIATED CONTENT

Supporting Information

The Supporting Information is available free of charge on the ACS Publications website at DOI: 10.1021/acs.chemma-ter.6b03221.

Explicit values of "design metrics" for materials screening; calculated exciton Bohr radii results; candidate compounds having slightly indirect band gaps; additional crystal orbital overlap populations analysis results; effect of spin–orbit coupling on band structures; decomposition enthalpy calculations with small molecules aligned along the $\langle 100 \rangle$ direction; evaluation of the steric sizes of organic molecular cations (PDF)

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Notes

The authors declare no competing financial interest.

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