## PHYS 5260: Quantum Mechanics - II

## Homework Set 1

Issued January 11, 2016 Due January 25, 2016

Reading Assignment: Shankar, Ch.16

1. Harmonic oscillator

Use a variational theory with a Gaussian trial wavefunction  $\psi(x) = Ne^{-\alpha x^2}$  (with N as proper normalization) to determine the ground state wavefunction, i.e., the optimum  $\alpha_0$ , and the corresponding ground state energy  $E_0$  by minimizing  $E(\alpha) = \langle H \rangle$  (where H is the harmonic oscillator Hamiltonian) with respect to the variational parameter  $\alpha$ . Compare your answers here with the exact results for  $\alpha_0$  and  $E_0$ .

- 2. (Shankar 16.1.3) For the attractive delta function potential  $V = -aV_0\delta(x)$ , use the variational principle with a Gaussian trial wavefunction to calculate the upper bound on  $E_0$  and compare it to the exact answer  $-ma^2V_0^2/2\hbar^2$  (from last semester, a problem that you should review).
- 3. Use the variational method with a variational function  $\psi(x) = Nxe^{-ax}$  (with N as proper normalization) for a 1D particle of mass m in a potential V(x) to determine the optimum value  $a_0$  of parameter a and the corresponding ground-state energy  $E_0 = E_{gs}(a_0)$ .

Take  $V(x) = \infty$ , for x < 0 and  $V(x) = \epsilon x$ , for 0 < x.

4. Helium atom within Hartree-variational approximation

Fill in all the steps for the variational estimate (with Z as the variational parameter) of the ground state energy of a Helium atom sketched out in Shankar, Ch.16, that leads to  $E_0$  in Eq.16.1.16.

5. Estimate the tunneling-out probability (decay rate) at energy E for the potential  $V(x) = \epsilon x$ , for 0 < x < d, V(x) = 0, for d < x, and  $V(x < 0) = \infty$ , using WKB approximation.

6. Use WKB approximation to compute a decay rate out of a 3D attractive short-ranged potential well for a particle in a state of orbital angular momentum  $\ell$  and energy E. How does the result depend on E and  $\ell$  in the universal low E limit?

Hint: Recall from last semester that the finite orbital angular momentum creates an additional effective centrifugal potential that qualitatively modifies the attractive short-ranged potential by introducing a repulsive power-law tail at large r. At low Eyou should be able to considerably simplify your computation.