

How do gravity and ice temperature affect the performance of thermal melting probes?

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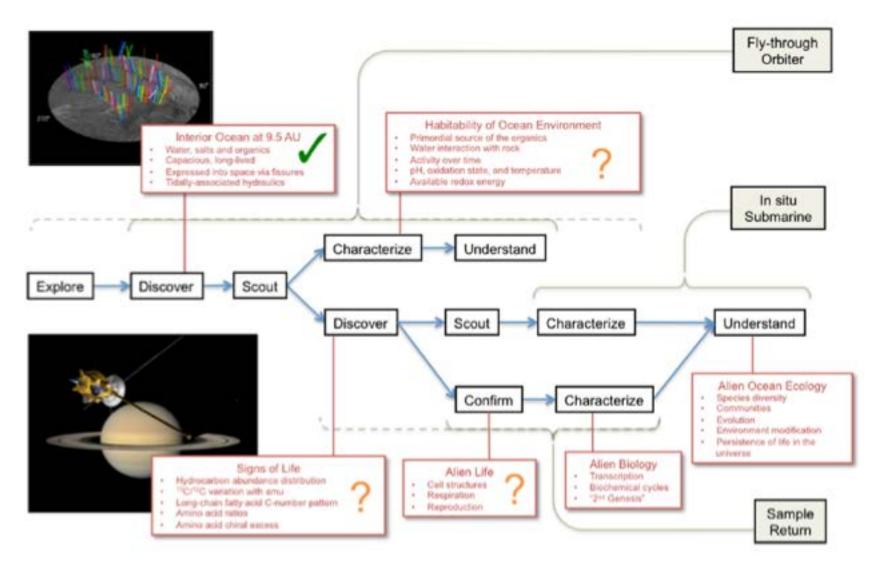
Smead Aerospace

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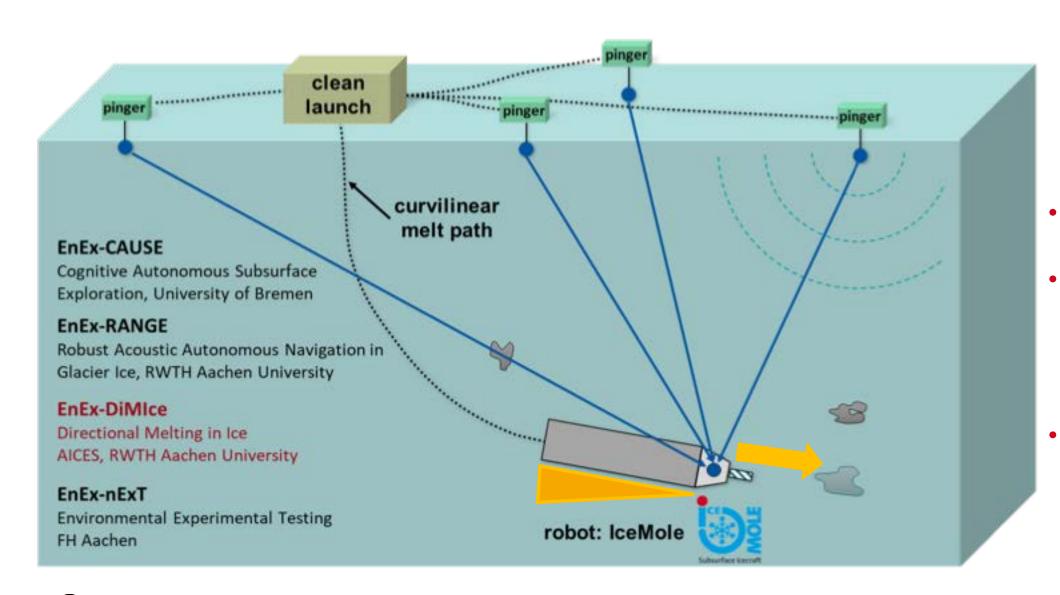
A roadmap to exploring the Ocean World Enceladus



Sherwood, Strategic map for exploring the ocean-world Enceladus. Acta Astronautica, 126: 52/58, 2016.



DLR Enceladus Explorer





- **before 2012** robotic student project
 - 2012 2015 **EnEx-MIDGE collaboration** advancing technologies and subglacial sample return in Antarctica
- 2015

further development of key technologies





on the basis of a decision by the German Bundestag



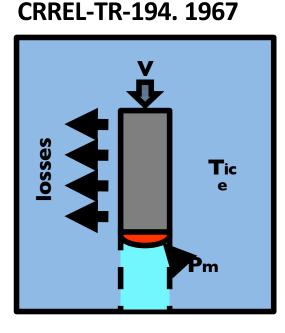


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Modeling the dynamics of melting probes – the 0D approach

Engineering model:

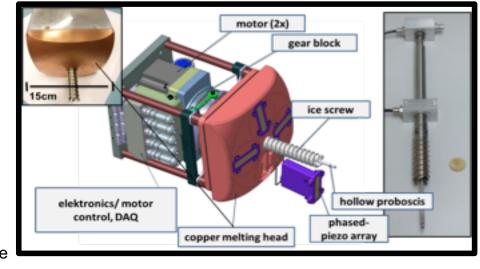
energy balance yields:



 $\begin{array}{l} \mbox{melting}\\ \mbox{velocity} \end{array} = \frac{\mbox{input power}}{\mbox{energy needed}} \\ V = \frac{P_m}{A\rho \left(h_m + c_p (T_m - T_{\rm ice})\right)} \end{array}$

V: melting velocity
P_m: input power
A: crosssection of the probe
ρ: density of the ice

c_p: specific heat capacity of the ice
 T_m: melting temperature
 T_{ice}: ice temperature
 h_m: melting enthalpy of ice



The IceMole's 2015 design:

How do melting probes perform in low gravity conditions?





Aamot,

Modeling the dynamics of melting probes – the 4D approach

Robot motion (concentrated)

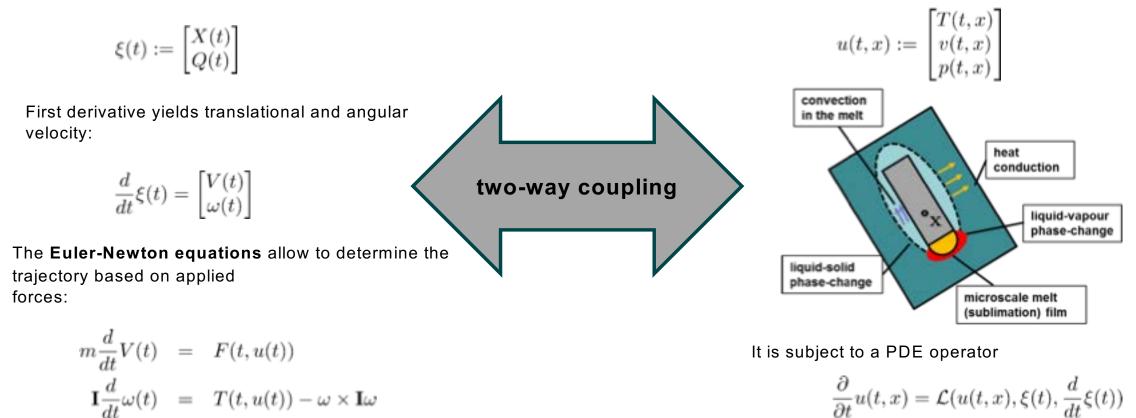
The current state of the probe is given by its center-of-mass and its attitude:

Cryoenvironment (distributed)

The current ambient state of the ambient is given by temperature, velocity and pressure:

that depends on position and attitude of the

probe, as well as its melting velocity.



The forces depend on the position of the liquid-solid

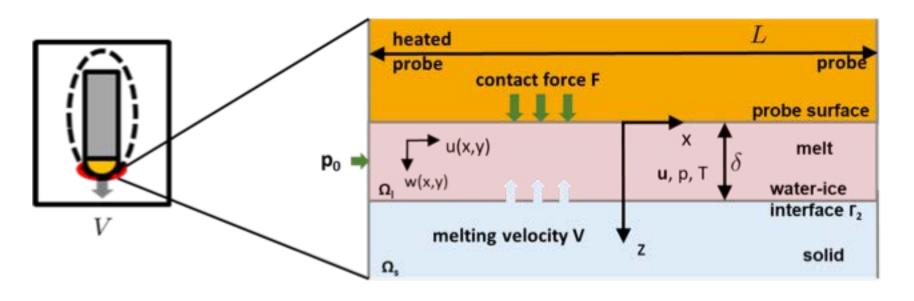
interface, hence the ambient state *u*.





Modeling the dynamics of melting probes – the smart way

Microscale melt film determines the probe's macroscale dynamics



Ω: Mass, momentum and energy balance:

$$\nabla \cdot \mathbf{u} = 0$$

$$\partial_t \mathbf{u} + (\mathbf{u} \cdot \nabla) \mathbf{u} = -\frac{1}{\rho} \nabla p + \nu \Delta \mathbf{u}$$

$$\partial_t T_l + (\mathbf{u} \cdot \nabla) T_l = \alpha_l \Delta T_l$$

Model reduction based on scaling arguments

Water-ice interface conditions:

- no-slip
- inflow according to melting rate
- melting temperature
- Stefan condition

Heat source surface:

• no-slip

 Ω_s : Heat equation in the solid ice

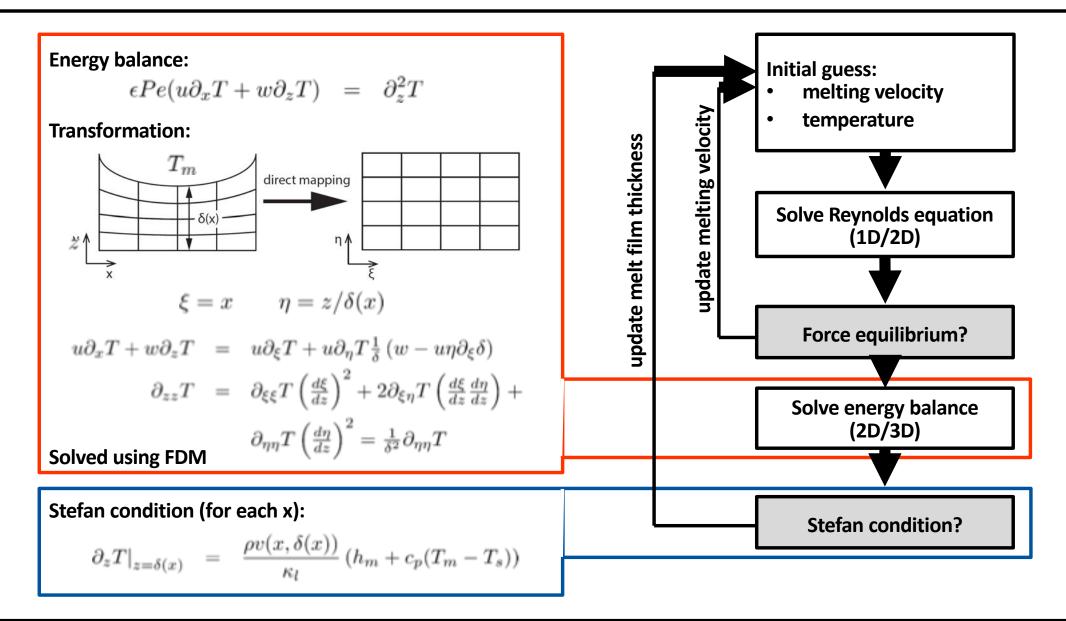
 $\partial_t T - (\mathbf{V} \cdot \nabla) T_s = \alpha_s \triangle T_s$

- no inflow
- temperature or heat flux
- Newton's third law:

$$\oint_{\text{surface}} p \, d\sigma = F$$



Hybridized computational model: SimCoMet – Simulating Contact Melting





Contact melting – some fundamental results

Rotational melting modes



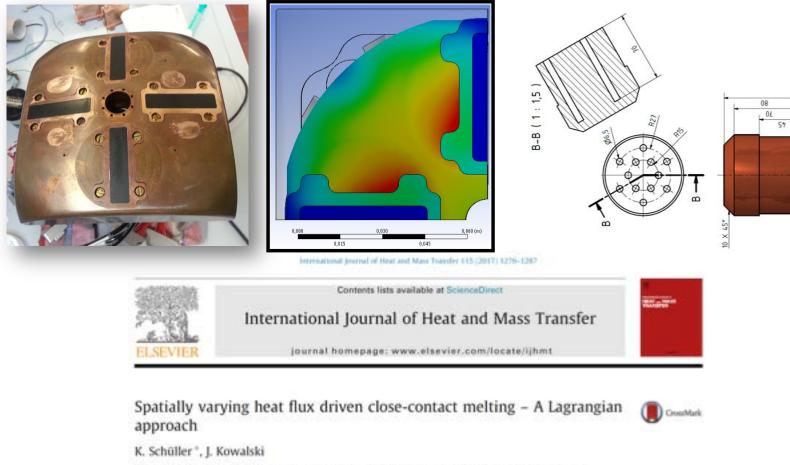
Contents lats available at Belandforent International Journal of Heat and Mass Transfer

Curvilinear melting - A preliminary experimental and numerical study

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Spatially varying heat flux distribution

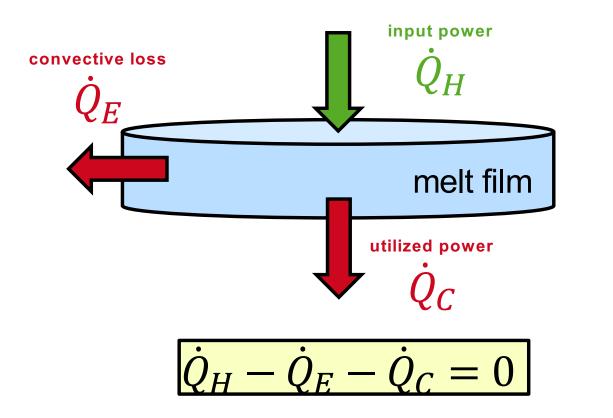


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Consider melt film (red) as a closed system:



and apply contact melting theory

Allows us to

- study the melting velocity / efficiency
- determine the critical refreezing length



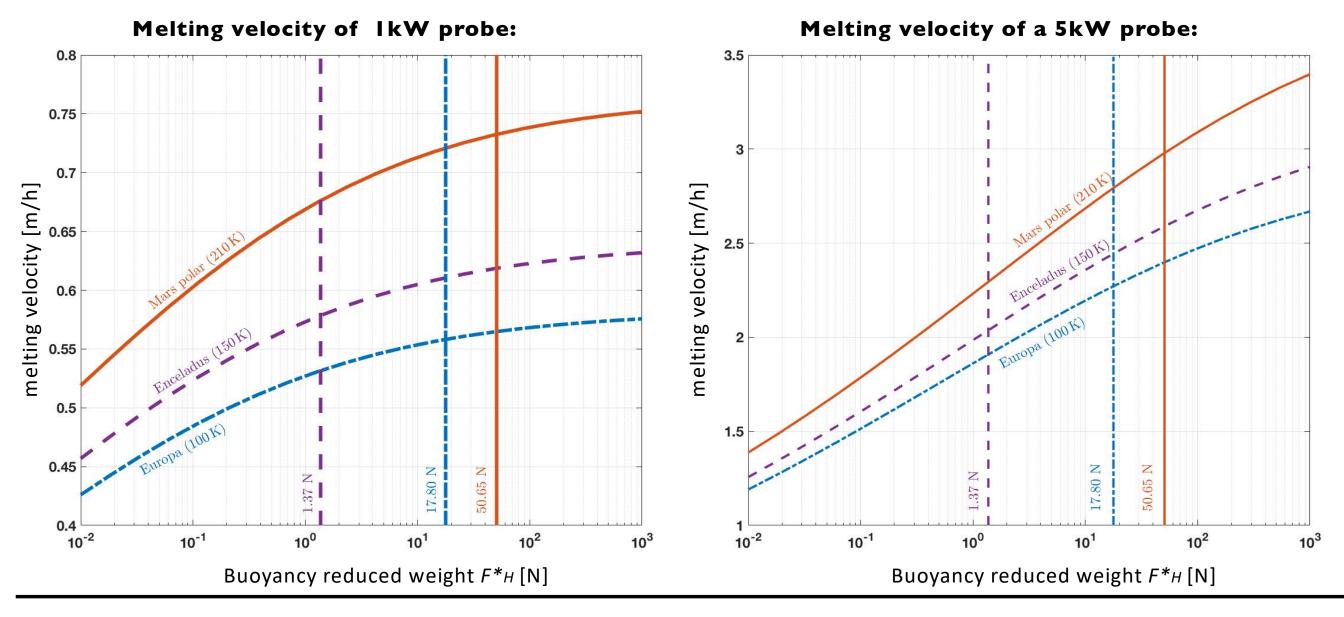
for a **reference probe** (1m long / 0.06m radius / 25kg) in a **representative cryoenvironments**:

- Mars: 210 K / 3.7 m/s²
- Enceladus: 150 K / 0.1 m/s²
- Europa: 100 K / 1.3 m/s²



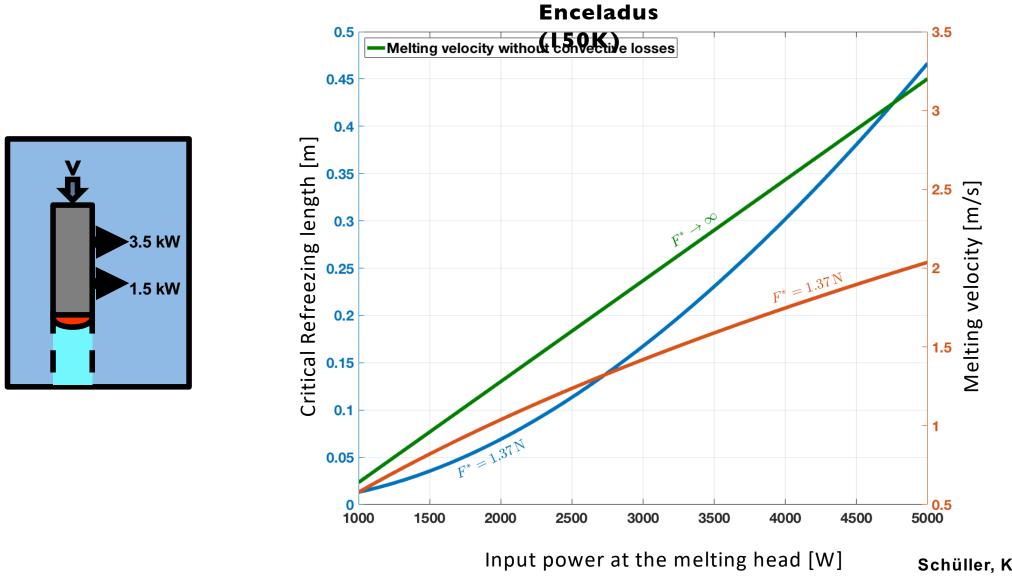


Melting velocity over contact force





Implications for the melting probe's design



Schüller, Kowalski, Icarus (accepted)



Conclusions

- We developed a flexible micro-scale contact melting simulation model
- We gained further insight into the behavior of melting probes in extreme cryoenvironments
- We contributed to the fundamental understanding of contact melting processes
- First validation experiments have been promising

Next steps

- Trajectory model for the IceMole (trajectory control will be tested in 2018 field campaign)
- High altitude rocket and vacuum chamber experiments due soon (VIPER / EnEx- nExT)
- In proposal phase: Smart process model and data integration for ice exploration











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