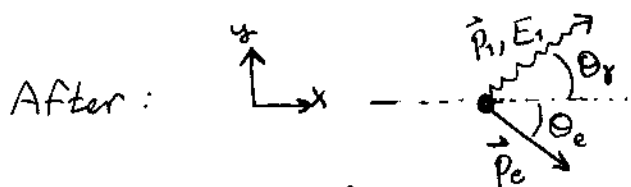


P3320 LECTURE 39

A more complicated example: Compton scattering. A γ -ray (massless, so $p^M p_M = 0$) scatters off of an atomic electron. The photon energy is huge compared to the electron's kinetic energy, so we can assume the electron is at rest.

Before: \vec{p}_0, E_0 \rightarrow \bar{e} at rest



where angles are measured relative to initial γ momentum.

We set the axes so interaction plane is xy .

What is E_1 as a function of θ_γ ?

Write out initial 4-momentum: ($m_e \equiv$ electron mass)

$$P_{\text{init}}^M = \left(\frac{E_0}{c}, p_0, 0, 0 \right) + (m_e c, 0, 0, 0)$$

$$P_{\text{final}}^M = \left(\frac{E_1}{c}, p_1 \cos \theta_\gamma, p_1 \sin \theta_\gamma, 0 \right) + \left(\frac{E_e}{c}, p_e \cos \theta_e, -p_e \sin \theta_e, 0 \right)$$

Know that each component $P_{\text{init}}^M = P_{\text{final}}^M$, so:

$$\left. \begin{aligned} \frac{E_0}{c} + \frac{m_e c^2}{c} &= \frac{E_1}{c} + \frac{E_e}{c} \\ \text{or } E_0 + m_e c^2 &= E_1 + \sqrt{m_e^2 c^4 + p_e^2 c^2} \end{aligned} \right\} \text{Total energy conservation}$$

$$p_0 = p_1 \cos \theta_\gamma + p_e \cos \theta_e \quad \left. \vphantom{p_0} \right\} p^1 \text{ conserved}$$

$$0 = p_1 \sin \theta_\gamma - p_e \sin \theta_e \quad \left. \vphantom{0} \right\} p^2 \text{ conserved.}$$

First, use p^2 equation to eliminate θ_e : $\sin \theta_e = \frac{p_1}{p_e} \sin \theta_\gamma$

But, since photons are massless, $E = pc$ always, so:

$$\sin \theta_e = \frac{E_1}{p_e c} \sin \theta_\gamma \quad \text{and we can eliminate } p_0, p_1.$$

substitute for $\cos\theta_e$:

$$\text{Now, } p_0 = \frac{E_0}{c} = \frac{E_1}{c} \cos\theta_1 + p_e \sqrt{1 - \left(\frac{E_1}{p_e c} \sin\theta_1\right)^2} \quad (*)$$

Combine with the 0th component equation

$$E_0 + m_e c^2 = E_1 + \sqrt{m_e^2 c^4 + p_e^2 c^2}$$

$$(E_0 - E_1 + m_e c^2) = \sqrt{m_e^2 c^4 + p_e^2 c^2}$$

Eliminate p_e in (*), and after lots of messy, nasty, ugly algebra:

$$E_1 = \frac{E_0}{1 + \frac{E_0}{m_e c^2} (1 - \cos\theta)}$$

→ Compton scattering is a dramatic confirmation of both special relativity and the particle nature of high-energy photons.

Forces and acceleration:

Know from problem set that $\vec{F} = m\vec{a}$ is not Lorentz invariant. Actually, it isn't even right (unless you use relativistic mass). Written as

$$\vec{F} = \frac{d\vec{p}}{dt} = \gamma m \frac{d\vec{u}}{dt}$$

it is correct, but note that it's still not Lorentz invariant, nor is it even a good four-vector.

Can create a four-force (Minkowski force) $K^M = \frac{dp^M}{d\tau}$ where $K^i = \gamma F^i$. However, when we calculate fields in the lab frame we end up more naturally with \vec{F} than K^M .

Note that the work-energy theorem still applies:

$$W = \int d\vec{l} \cdot \vec{F} = E_{\text{final}} - E_{\text{initial}} \quad \text{where } E = \sqrt{p^2 c^2 + m^2 c^4}$$

One last mechanics/transformation item: velocity addition rule:

Take a particle moving at u in lab frame, and boost it to a frame moving at v : what is its velocity in primed frame $\frac{dx'}{dt'}$?

$$u = \frac{dx}{dt} \quad \text{but } dx' = \gamma(dx - v dt)$$

$$dt' = \gamma\left(dt - \frac{v}{c^2} dx\right)$$

$$u' = \frac{dx'}{dt'} = \frac{dx - v dt}{dt - \frac{v}{c^2} dx} = \frac{\frac{dx}{dt} - v}{1 - \frac{v}{c^2} \frac{dx}{dt}} = \boxed{\frac{u - v}{1 - \frac{uv}{c^2}}}$$

(Einstein's velocity addition rule)