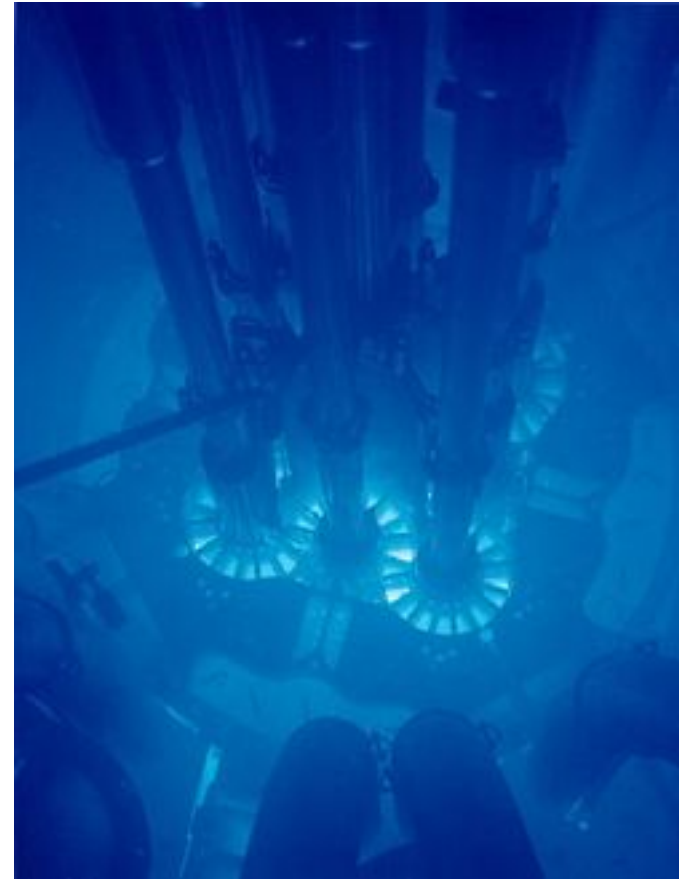


# Special relativity

## Announcements:

- Homework solutions will be available around 4pm today.
- Next weeks homework should be available by 7pm today.
- The first midterm exam is on February 17 (two weeks from yesterday)



Today we will pretty much finish up special relativity.

# Relativistic momentum and energy

Relativistic momentum:  $\vec{p} = \gamma_u m \vec{u}$

Total energy:  $E = \gamma_u mc^2$

With these definitions for momentum and energy, conservation of momentum and conservation of (total) energy continue to work in isolated relativistic systems

At rest,  $\gamma_u = 1$  so the rest energy is  $E_{\text{rest}} = mc^2$

This tells us that mass and energy are equivalent. Mass is a type of energy.

Kinetic energy is *always* the total energy minus the rest energy:

$$KE = E - E_{\text{rest}} = \gamma_u mc^2 - mc^2 = (\gamma_u - 1)mc^2$$

At *low* speeds, this becomes the familiar  $KE \approx \frac{1}{2} mu^2$

At what speed is the total energy of a particle equal to twice its rest mass energy?  $E = \gamma_u mc^2$      $E_{\text{rest}} = mc^2$      $KE = (\gamma_u - 1)mc^2$

- A. 0
- B. 0.7c
- C. 0.87c**
- D. 0.94c
- E. c

To have total energy equal to twice the rest mass energy, need  $\gamma=2$

Solve  $\gamma = (1 - \beta^2)^{-1/2}$  for  $\beta$ .

$$\gamma^2 = \frac{1}{1 - \beta^2} \longrightarrow \frac{1}{\gamma^2} = 1 - \beta^2 \longrightarrow \beta^2 = 1 - \frac{1}{\gamma^2}$$

$$\beta = \sqrt{1 - \frac{1}{\gamma^2}} = \sqrt{1 - \frac{1}{2^2}} = \sqrt{1 - 0.25} = \sqrt{0.75} = 0.87$$

so you need to be moving pretty fast to get your kinetic energy close to your rest mass energy!

# Relativistic momentum and energy

$$\text{Relativistic momentum: } \vec{p} = \gamma_u m \vec{u}$$

$$\text{Total energy: } E = \gamma_u mc^2$$

With some algebra we can eliminate the velocity variable from these two relations and obtain the triangle relation:

$$E^2 = (mc^2)^2 + (pc)^2$$

A new unit of energy is the electron-volt (eV). It's the energy obtained by an electron moving through 1 V. It is not an SI unit but is very common.

$$\Delta E = q\Delta V = 1 \text{ eV} = 1.6 \cdot 10^{-19} \text{ C} \cdot 1 \text{ V} = 1.6 \cdot 10^{-19} \text{ J}$$

Since  $mc^2$  is a unit of energy, dividing energy by  $c^2$  gives a unit of mass. Also, dividing energy by  $c$  gives a unit of momentum.

A proton has a mass of  $938 \text{ MeV}/c^2$ . What is this in kg?

$$938 \text{ MeV} / c^2 = \frac{938 \times 10^6 \text{ eV} \cdot 1.60 \times 10^{-19} \text{ J/eV}}{(3.00 \times 10^8 \text{ m/s})^2} = 1.67 \times 10^{-27} \text{ kg}$$

Also use eV/c or MeV/c units for momentum

Please answer this question on your own.  
No discussion until after.

True or false: A massless particle has a finite momentum.

A. True

B. False

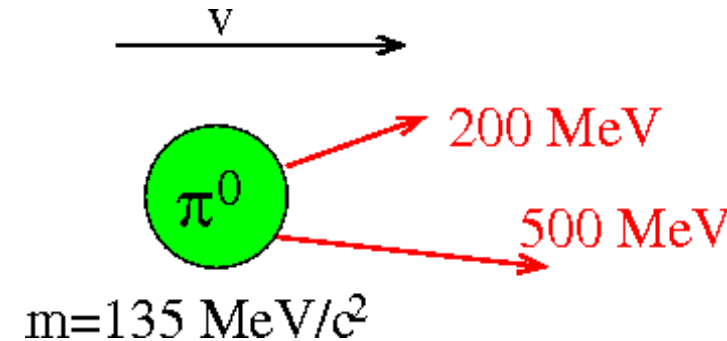
C. More than one of the above

D. None of the above

We will see how in the next few minutes

# Conservation of energy for subatomic particles

A  $\pi^0$  particle decays into two photons. The photon energies are measured to be 200 MeV and 500 MeV. The  $\pi^0$  rest mass is  $135 \text{ MeV}/c^2$ . What is the  $\pi^0$  velocity?



Conservation of energy requires that the sum of the photon energies equal the total  $\pi^0$  energy.

$$E(\pi^0) = E(\gamma_1) + E(\gamma_2) = 200 \text{ MeV} + 500 \text{ MeV} = 700 \text{ MeV}$$

Although we don't need it, we can see that the  $\pi^0$  kinetic energy is  $KE = E - mc^2 = 700 \text{ MeV} - 135 \text{ MeV} = 565 \text{ MeV}$

Can get velocity from

$$\gamma = \frac{1}{\sqrt{1 - \beta^2}}$$

From  $E = \gamma mc^2$

we note that  $\gamma = \frac{E}{mc^2}$

# Conservation of energy for subatomic particles

What is the velocity of a  $\pi^0$  with rest mass  $135 \text{ MeV}/c^2$  and total energy  $700 \text{ MeV}$ ?

$$\gamma = \frac{E}{mc^2} = \frac{700 \text{ MeV}}{135 \text{ MeV} / c^2 \cdot c^2} = \frac{700 \text{ MeV}}{135 \text{ MeV}} = 5.2$$

We found before that we could solve  $\gamma = \frac{1}{\sqrt{1-\beta^2}}$

to get  $\beta = \sqrt{1 - \frac{1}{\gamma^2}}$

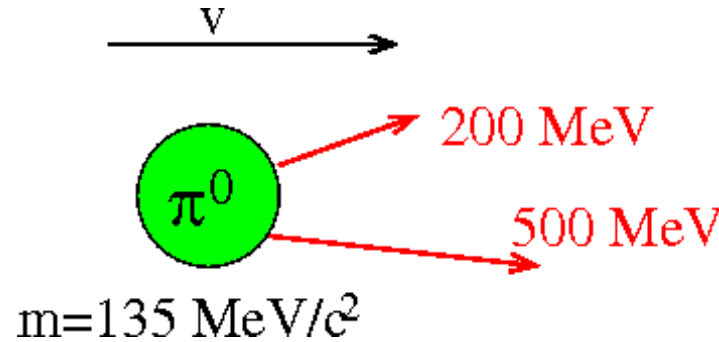
So this gives us:  $\beta = \sqrt{1 - \frac{1}{\gamma^2}} = \sqrt{1 - \frac{1}{5.2^2}} = \sqrt{0.96} = 0.98$

So the  $\pi^0$  velocity was  $0.98c$

What is the  $\pi^0$  momentum?

From the relativistic equation for momentum we get

$$p = \gamma mu = 5.2 \cdot 135 \text{ MeV} / c^2 \cdot 0.98c = 690 \text{ MeV} / c$$



Q. What is the net momentum of the two photons?

A. 0

**B. 690 MeV/c**

C. 700 MeV/c

D. 135 MeV/c

Momentum is conserved so the  $\pi^0$  momentum is carried by the photons

But photons are massless so  $p = \gamma mu$  doesn't work. What does?

# Energy and momentum of massive & massless particles

Remember our triangle relation:  $E^2 = (pc)^2 + (mc^2)^2$

For a massless particle like a photon, this reduces to  $E = |pc|$   
so the energy and momentum are the same

Keep in mind that momentum is a vector and energy is a scalar

As you will find out soon, Einstein won the Noble Prize for finding the energy of a photon is  $E = hf$

$f$  = frequency and  $h$  is Plank's constant:  $6.63 \times 10^{-34} \text{ J} \cdot \text{s} = 4.14 \times 10^{-15} \text{ eV} \cdot \text{s}$

This can then be used to obtain the photon momentum

Please answer this question on your own.  
No discussion until after.

The correct formula for Newton's 2<sup>nd</sup> law in special relativity is:

A.  $\vec{F} = \frac{Gm_1m_2}{r^2} \hat{r}$

B.  $\vec{F} = m\vec{a}$

C.  $\vec{F} = \gamma m\vec{a}$

D.  $\vec{F} = \frac{d\vec{p}}{dt}$

Forces still work in relativity, pretty much the same as in classical physics.

Since we need to allow for massless particles, we need to use  $\vec{F} = d\vec{p} / dt$  for Newton's 2<sup>nd</sup> law.

# The other “derivation” of $E=\gamma mc^2$

Back when we were discussing spacetime diagrams I mentioned that just like distance between two points is invariant in 3D, there is a 4D invariant called the *spacetime interval*

$$\Delta s^2 = (c\Delta t)^2 - (\Delta x)^2 - (\Delta y)^2 - (\Delta z)^2 = (c\Delta t)^2 - (\Delta \vec{r})^2$$

Or, back to just one space dimension:  $\Delta s^2 = (c\Delta t)^2 - (\Delta x)^2$

What happens if we multiple everything by  $\frac{m^2 c^2}{(\Delta \tau)^2}$  where  $\Delta \tau$  is the *proper time*

$$m^2 c^2 \left( \frac{\Delta s}{\Delta \tau} \right)^2 = m^2 c^4 \left( \frac{\Delta t}{\Delta \tau} \right)^2 - c^2 \left( m \frac{\Delta x}{\Delta \tau} \right)^2$$

This is  $\gamma$ 
This is  $\gamma m v = p$

This is  $c$

We end up with a different writing of the triangle equation:  
 $(mc^2)^2 = (\gamma mc^2)^2 - (pc)^2$   
 or

$$(mc^2)^2 = E^2 - (pc)^2$$

# Four vectors

Position and momentum are 3-vectors since in a 3D space, they measure three things  $\vec{r} = (x, y, z)$   $\vec{r} = (r, \theta, \phi)$   $\vec{p} = (p_x, p_y, p_z)$

In the 4 dimensions of spacetime we have 4-vectors which are a 3-vector plus a 4<sup>th</sup> scalar. Whether we put the 4<sup>th</sup> scalar at the end or beginning of the order list depends on the convention.

The 4<sup>th</sup> scalar associated with the position 3-vector is, of course, time or  $ct$ :  $(x, y, z, ct)$  or  $(ct, x, y, z)$  and the invariant spacetime interval is  $\Delta s^2 = (c\Delta t)^2 - (\Delta x)^2 - (\Delta y)^2 - (\Delta z)^2 = (c\Delta t)^2 - (\Delta \vec{r})^2$

The 4<sup>th</sup> scalar associated with the momentum 3-vector is energy or  $E/c$  to keep the units straight:  $(p_x, p_y, p_z, E/c)$  or  $(E/c, p_x, p_y, p_z)$  and the invariant is  $m^2 c^2 = \frac{E^2}{c^2} - p^2$  (our triangle relation again)

# End of special relativity

We will be starting something new on Friday. I haven't decided what yet. So no reading assignment for Friday.